

Seismic noises and earthquakes impact on the **Virgo** gravitational-wave detector during the **O3 run** (04/2019 – 03/2020)

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On behalf of the **Virgo Collaboration**
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Outline

- The O3 run in a nutshell
 - Virgo performance
- Seismic noise
 - Monitoring and disentangling its different contributions
 - Impact on the Virgo detector during the O3 run
- Earthquakes
 - Early warning system and control room mitigation
 - Virgo robustness against earthquakes
 - Control losses: from strong and distant earthquakes but also very close ones
- Prospects for the O4 run (2022-2023)

The O3 run

- Global network of ground-based gravitational-wave (GW) advanced detectors

- LIGO Hanford
- LIGO Livingston
- Virgo



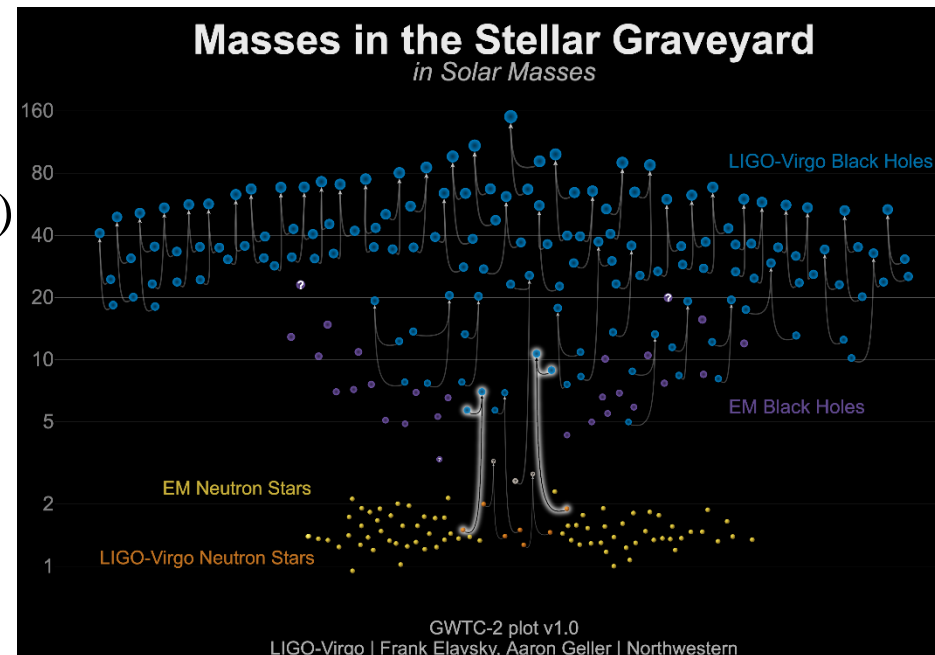
→ All three detectors taking data together from day 1 of the run

- O3: 2 data-taking periods

- O3a: 2019/04/01 → 2019/10/01
- O3b: 2019/11/01 → 2020/03/27
 - ◆ Premature ending due to covid-19
- 1-month commissioning break (10/2019)

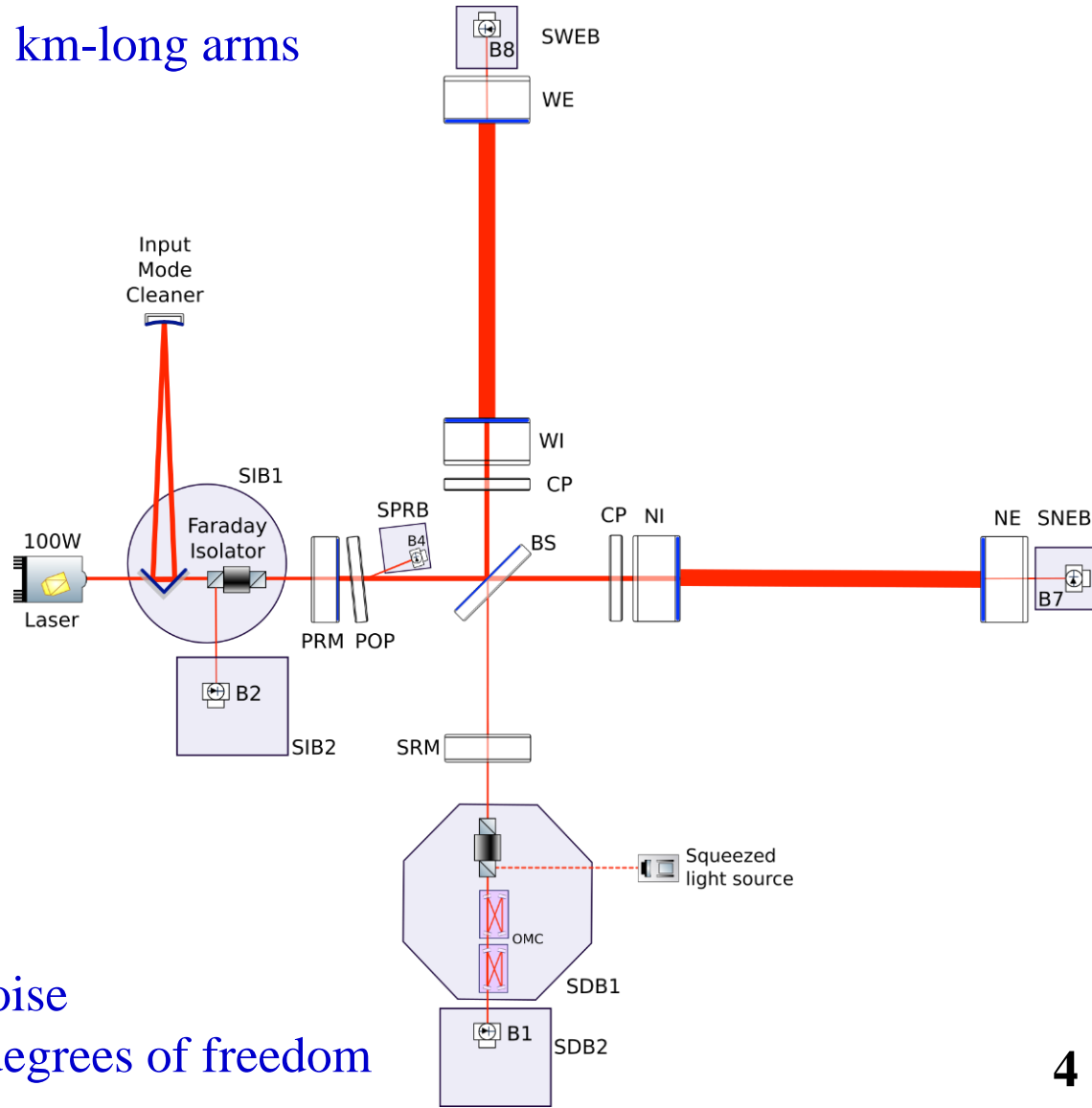
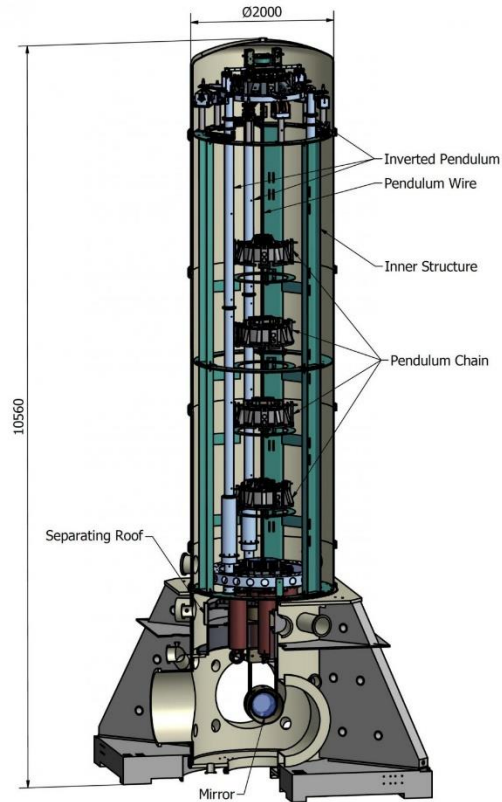
- Scientific harvest

- 56 public alerts
- Exceptional events
 - ◆ Neutron star – black hole mergers
- Updates to the GW Transient Catalog
 - ◆ Published for O3a – public data
 - ◆ In progress for O3b



The O3 run for the Virgo detector

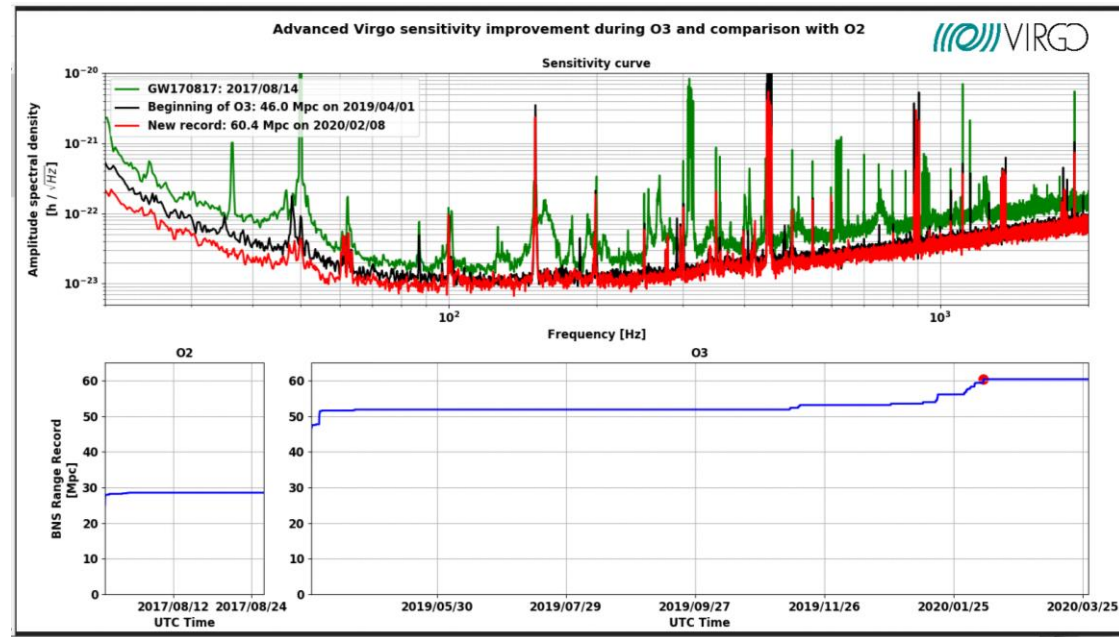
- **Power-recycled, suspended Michelson interferometer** with **Fabry-Perot cavities** in its **3 km-long arms**



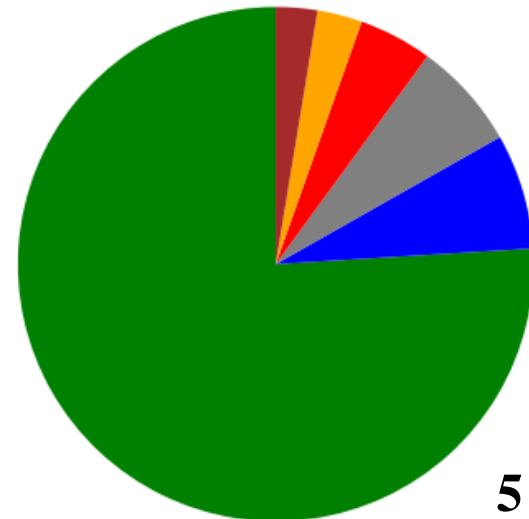
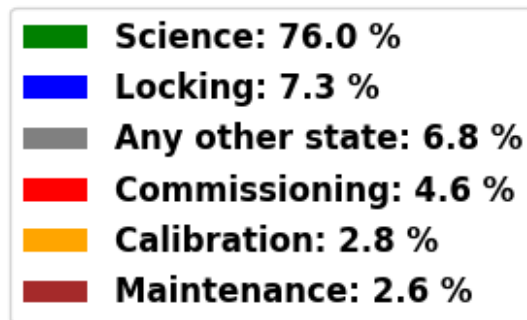
- **7 m-tall “superattenuators”** to
 - **Isolate mirrors** from seismic noise
 - **Control their position** in all 6 degrees of freedom

The O3 run for the Virgo detector

- Sensitivity about doubled w.r.t. the O2 run (August 2017)
 - Binary neutron star range (BNS range)
 - ◆ Steady: 45-55 Mpc
 - ◆ Record: 60 Mpc



- Duty cycle (O3a + O3b)
 - Science data taking: 76%
 - ◆ Consistent with O2 (80%)
 - Remaining time divided almost equally among 3 categories
 - ◆ Working point (re)acquisition
 - ◆ Maintenance + calibration + commissioning
 - ◆ Problems preventing the normal running of Virgo

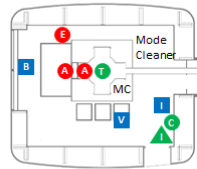


Environmental monitoring and seismic noise

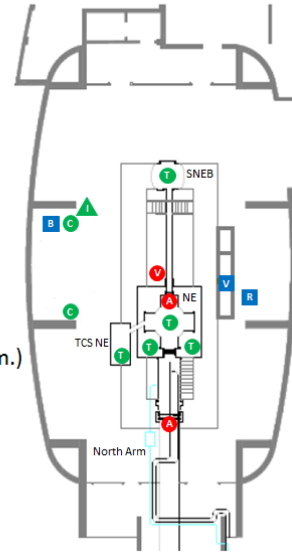
- **Virgo environmental sensor array**

- **MCB**: Mode-Cleaner Building
- **N(W)EB**: North (West) End Building
- **CEB**: Central Building

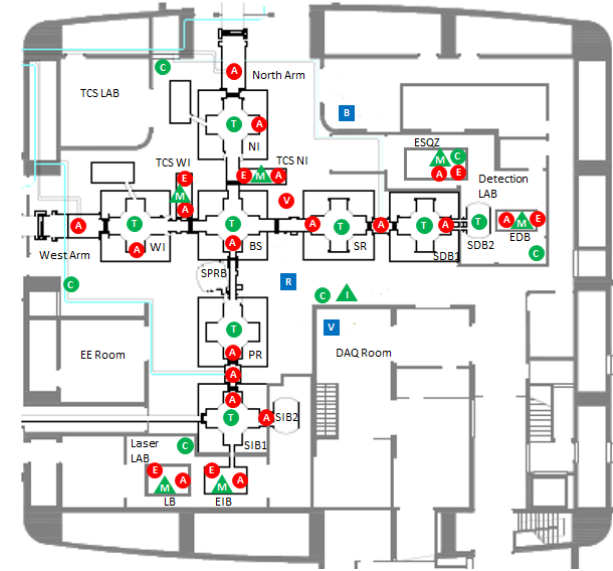
MCB



NEW (WEB)



CEB

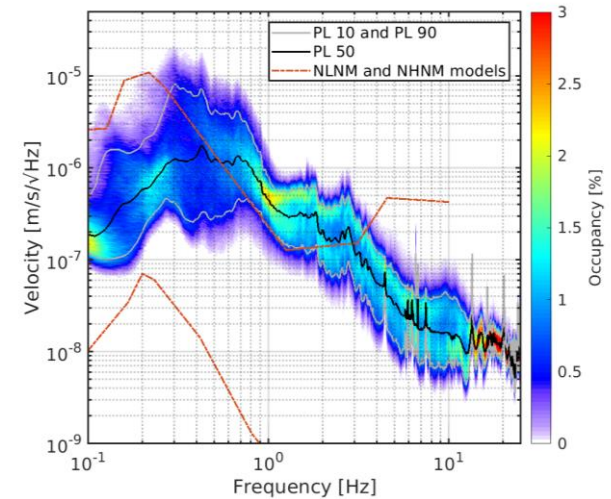


- Accelerometer
- Episensor
- Velocimeter
- Thermometer
- Comb. (temp.+press.+hum.)
- Microphone
- Infrasound microphone
- Magnetometer
- Voltage probe
- Current probe
- Radio frequency antenna

- **Seismic noise**

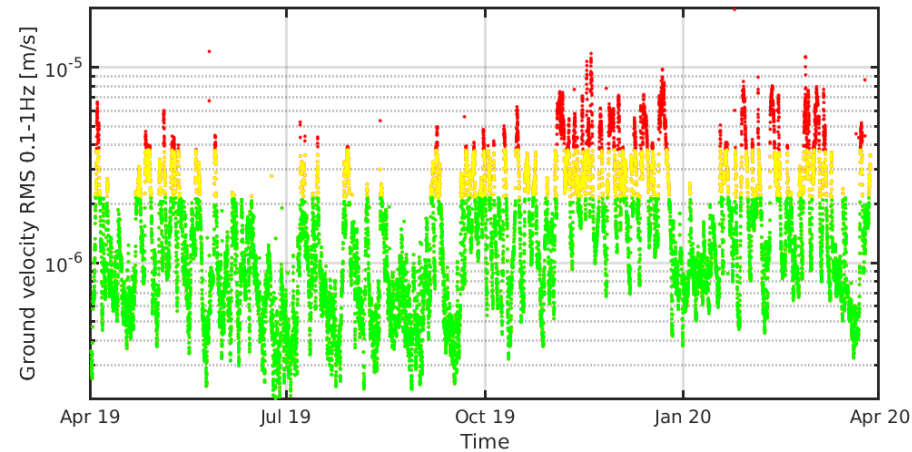
- **Microseism**: 0.1 ÷ 1 Hz
 - ◆ Dominant
 - ◆ Interaction between sea waves and ground
 - ◆ Peak around 350 mHz
- **Anthropogenic**: 1 ÷ 5-10 Hz
 - ◆ Heavy vehicles on elevated roads
- **Onsite**: 10 ÷ 40 Hz
 - ◆ Traffic on nearby roads, agricultural activities

→ **Frequency band-limited RMS (BLRMS)** to isolate the different contributions

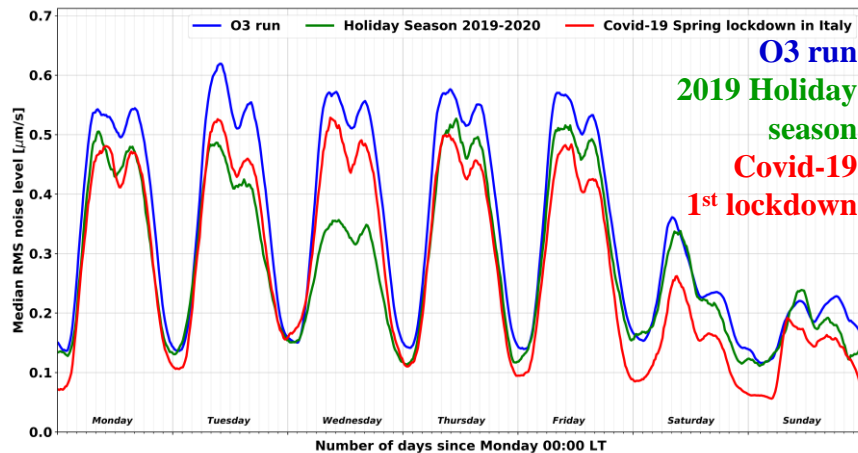


Seismic noise variability

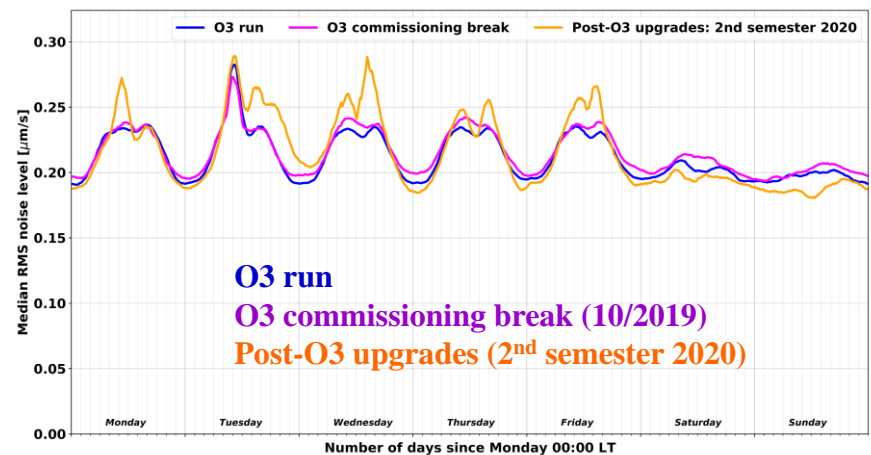
- **Microseism: seasonal variations**
 - Larger in Fall/Winter
 - Color code
 - ◆ **Green:** < 75th percentile
 - ◆ **Yellow:** 75th – 90th percentile
 - ◆ **Red:** > 90th percentile
- **Anthropogenic + on-site**
 - Impact of “global conditions”
 - ◆ Day/night + weekday variations
 - ◆ Holidays, pandemic...



Anthropogenic



On-site

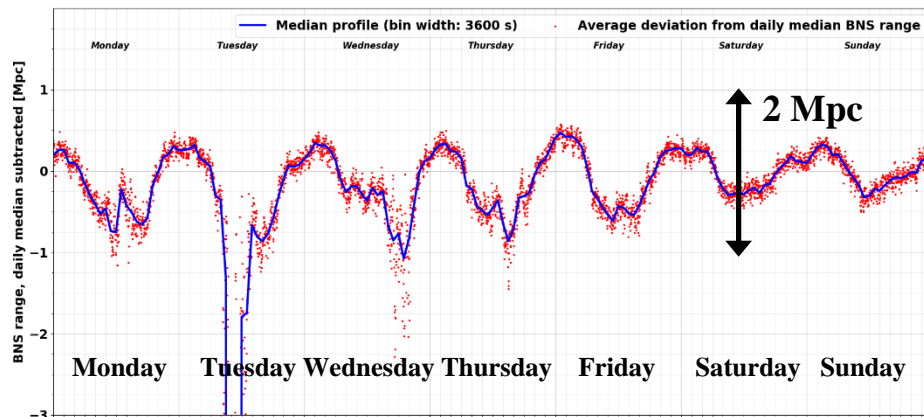


Sensitivity modulation

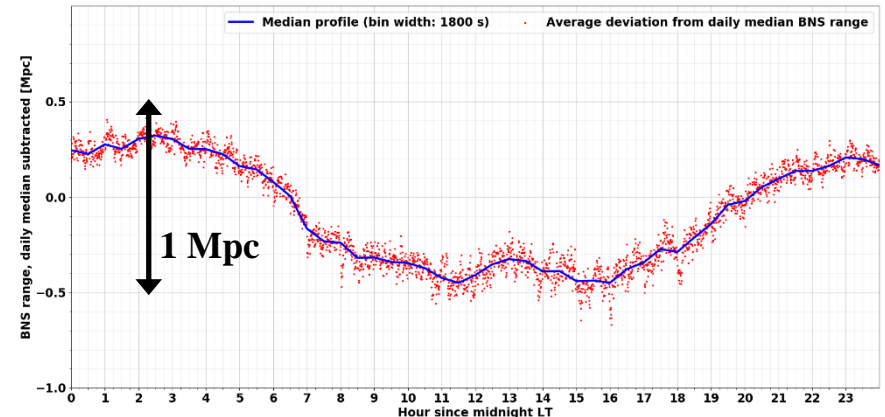
- Figure-of-merit: the **BNS range**
 - Subject to **variations from multiple origins** – not just the environment
 - ♦ Control accuracy, detector global status, minor problems, etc.
- Raw” BNS range value not suitable for such study
 - ♦ Instead: use **BNS range variations around its daily median level**

- **O3-averaged variations**

Over a week baseline



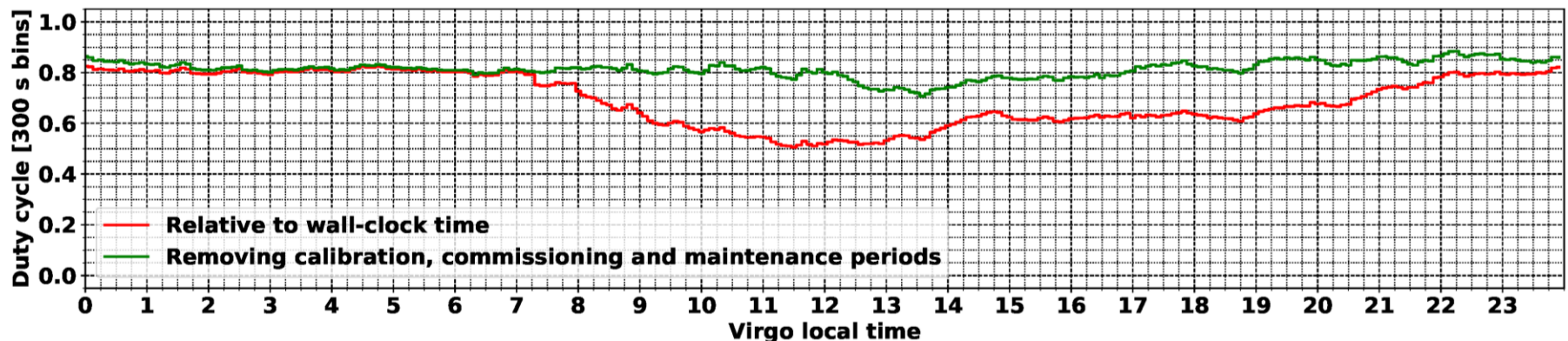
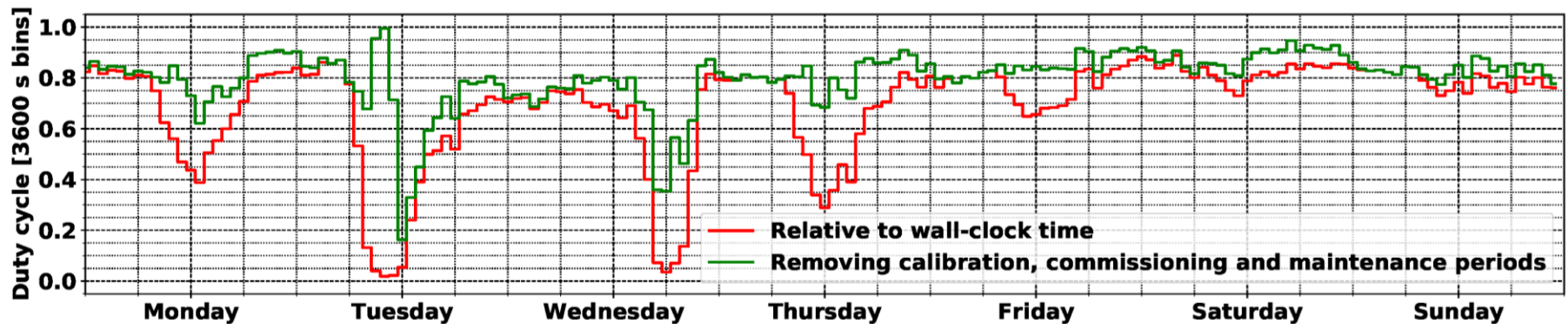
Over a 24-hour baseline



- **Modulation similar to anthropogenic noise**
 - **Limited amplitude: a few percents at most**

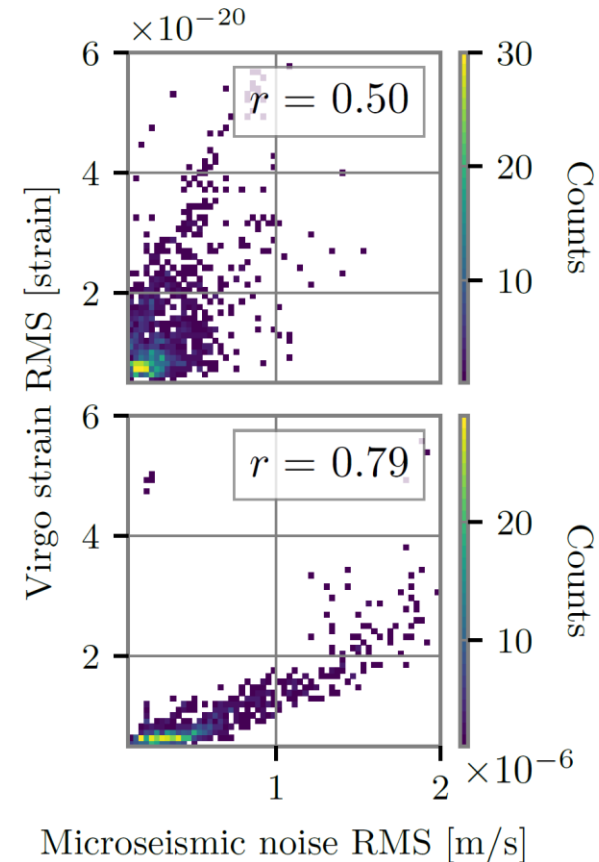
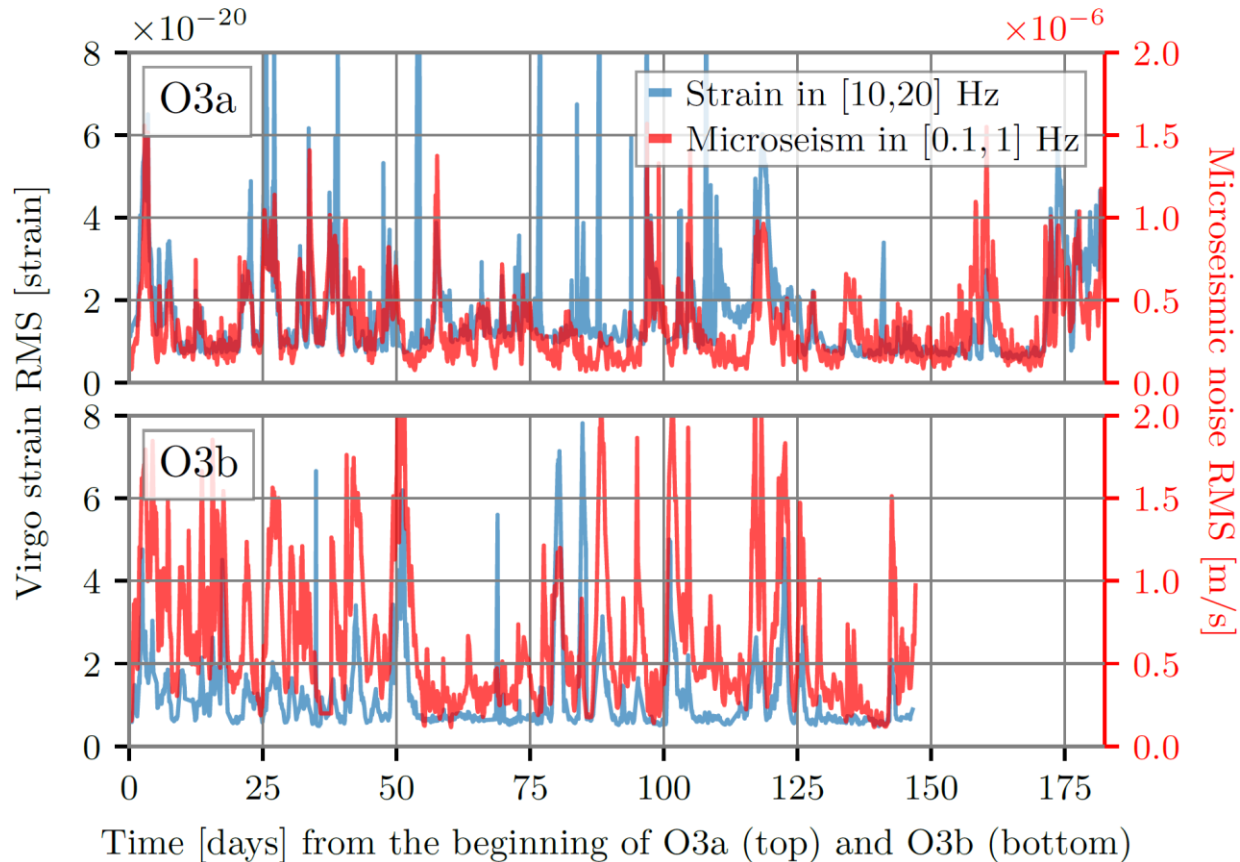
Duty cycle

- Defined as the **fraction of time spent taking Science data**
 - **Red**: relative to **wall-clock time**
 - **Green**: removing calibration, commissioning and maintenance periods
- **Not the same modulation**
 - Duty cycle driven by the (lack of) activities of the crew on duty



Microseism impact

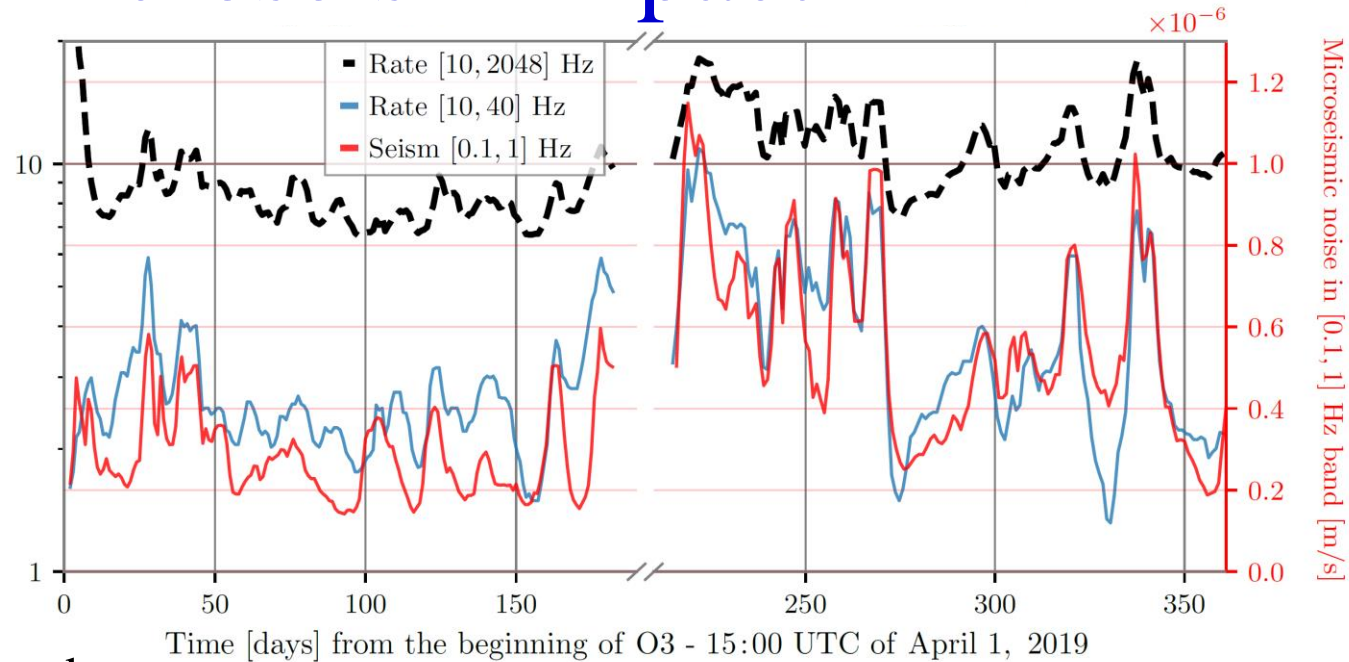
- Reconstructed **GW strain channel $h(t)$** worsens during high microseism activity
 - Up to ~ 40 Hz



- **Blue:** $h(t)$ BLRMS
- **Red:** Microseismic band BLRMS

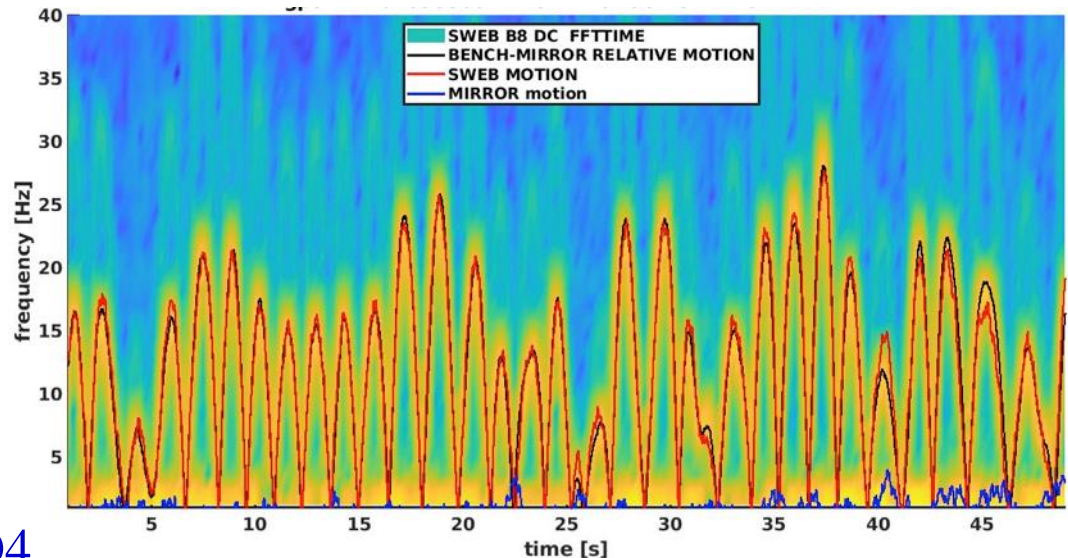
Microseism impact

- **Noise transient rate per minute**
 - Black: rate in the 10 ÷ 2048 Hz band
 - Blue: rate in the 10 ÷ 40 Hz band
 - Red: **microseism BLRMS**



→ **Impacts data quality** and **GW search trigger rate**

- **Main path identified**
 - High microseism
 - Larger relative motion of a detector component
 - **Scattered light**
 - Typical “arches” in spectrograms

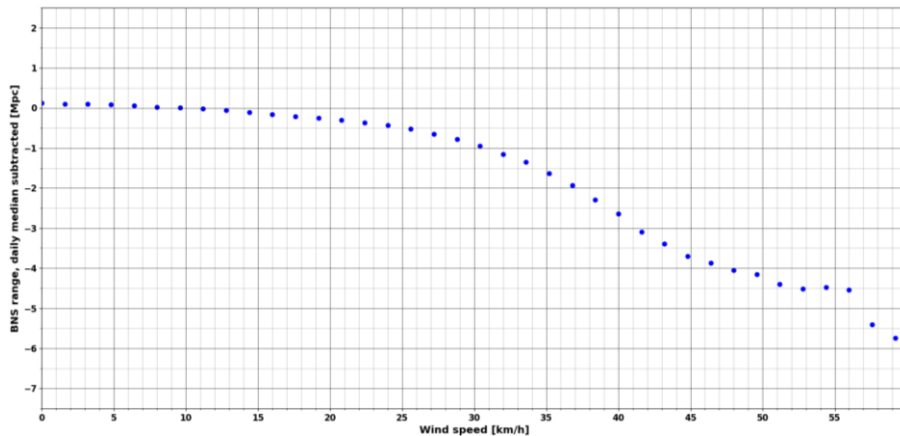


→ **Improvements foreseen for O4**

Microseismic noise in [0.1, 1] Hz band [m/s]

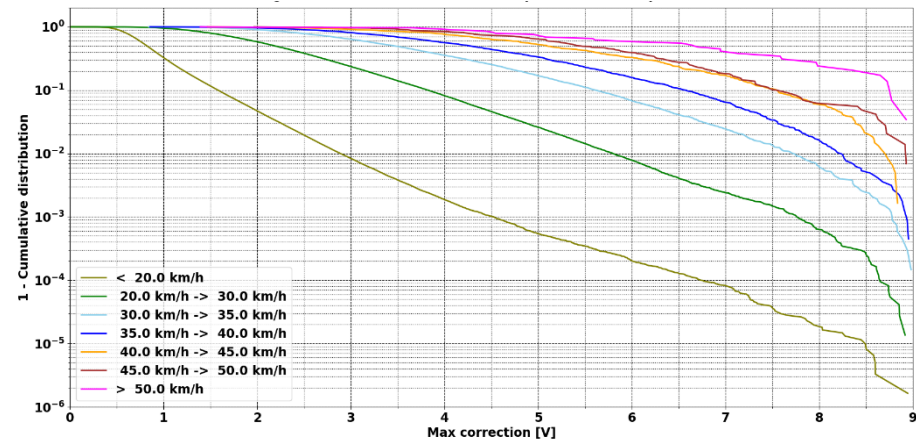
Wind impact

- **Bad weather** \Leftrightarrow high microseism activity (rough sea) and wind
→ **Disentangling the two contributions**
- **Some wind impact** on the BNS range above ~ 25 km/h



- **Up to 10% variation:**
significant but limited
- **Detector robustness**

- **Larger corrections** to keep the detector control as the wind speed increases



- **Limited actuation range**
- **Saturation:** immediate control loss

Wind impact

- Duty cycle

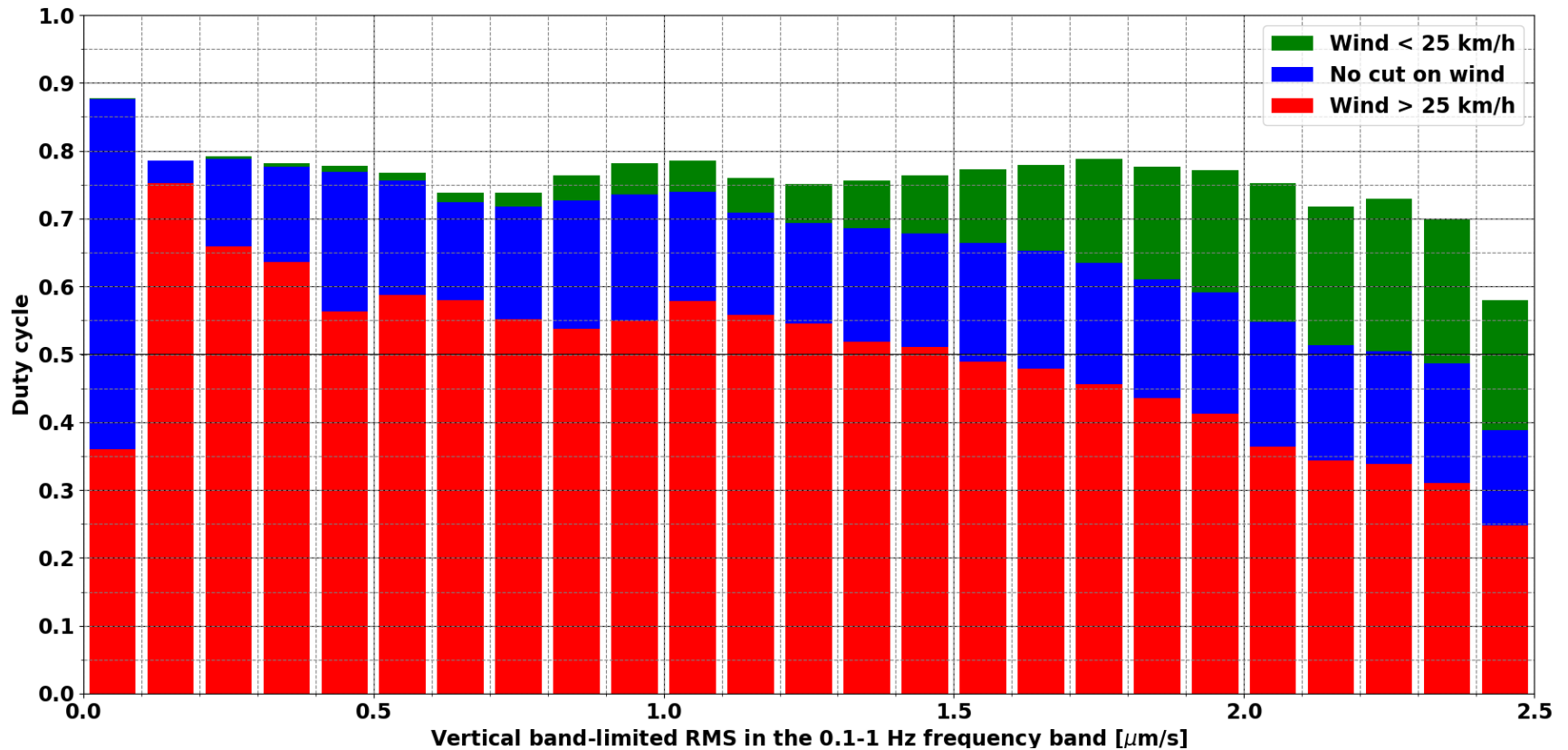
- x-axis: microseism BLRMS

- 3 datasets

- ◆ Blue: no cut on wind

- ◆ Green: low wind

- ◆ Red: high wind



→ Detector robust against microseism but more sensitive to wind

Earthquakes

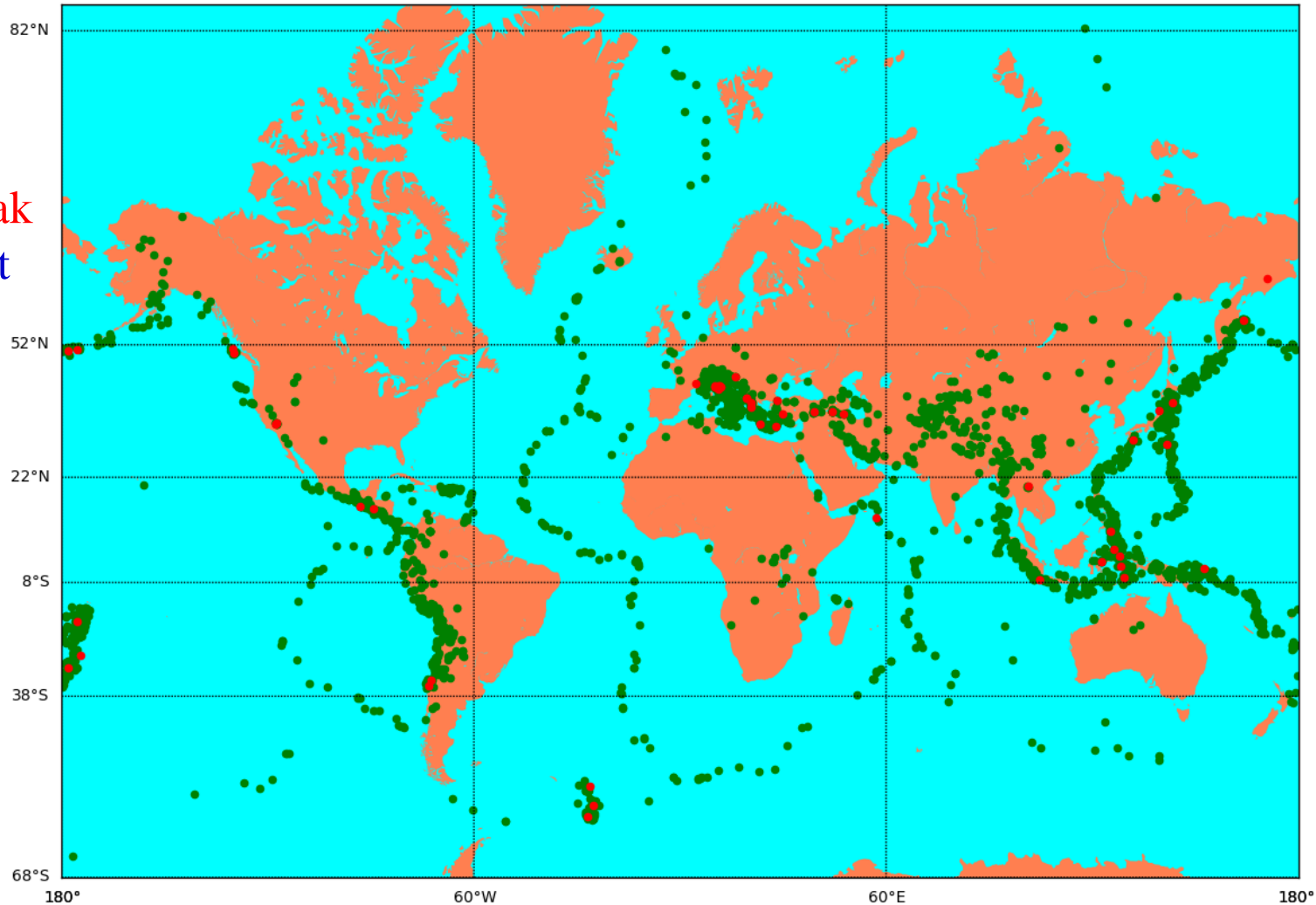
- High-enough seismic waves \Rightarrow feedback system saturation
 - \Rightarrow Working point not controlled anymore (“lock loss”)
 - \Rightarrow Duty cycle decreases
 - “Locking time” + eventually the time to damp excited suspensions
- Seismon: an earthquake early warning system
 - Developed by LIGO; running at EGO since O2
 - Input: earthquake alerts from a low-latency US Geological Survey (USGS) stream
 - Output: seismic wave arrival times and amplitude estimation at detector location
 - \rightarrow Interfaced with Virgo data acquisition and control system
- Earthquake mitigation
 - Requires warning to arrive in the control room prior to the seismic waves
 - ♦ Up to tens of minutes of margin for the most distant earthquakes on Earth
 - Manual switch to a more resilient control configuration w/o losing the lock
 - ♦ (Slightly) more noisy
 - ♦ Only validated for Science data taking close to the end of O3b
 - ♦ Actuation range doubled \Rightarrow saturation (and control loss) less likely
 - Back to nominal control when seismic waves fade away
 - ♦ Overall duty cycle gain if the detector has survived the earthquake

Control losses due to earthquakes

- **601 lock losses** from Science mode during the whole O3 run
 - Less than 2 / day in average
 - Locking phase median duration: **25 minutes**
 - ◆ Median number of attempts: **2**
- **30 (5%) found to be due to earthquakes**
 - About 1 / 10-11 days in average
 - 24 more lock losses due to earthquakes found while not taking Science mode data
 - **Included in the following analysis** to increase dataset studied
- **2 main categories**
 - **Distant** and **strong** earthquakes
 - ◆ Warning available ahead of the seismic waves but the control could not hold
 - **Weak** but **very close** earthquakes
 - ◆ Not reported at the output of Seismon
 - Found using the **Istituto Nazionale di Geofisica e Vulcanologia (INGV) public earthquake database**
 - ◆ Too close anyway to trigger “early” warnings
 - But important to **find the right cause for these lock losses**
 - ◆ Time-coincident USGS early warnings missing or not making sense

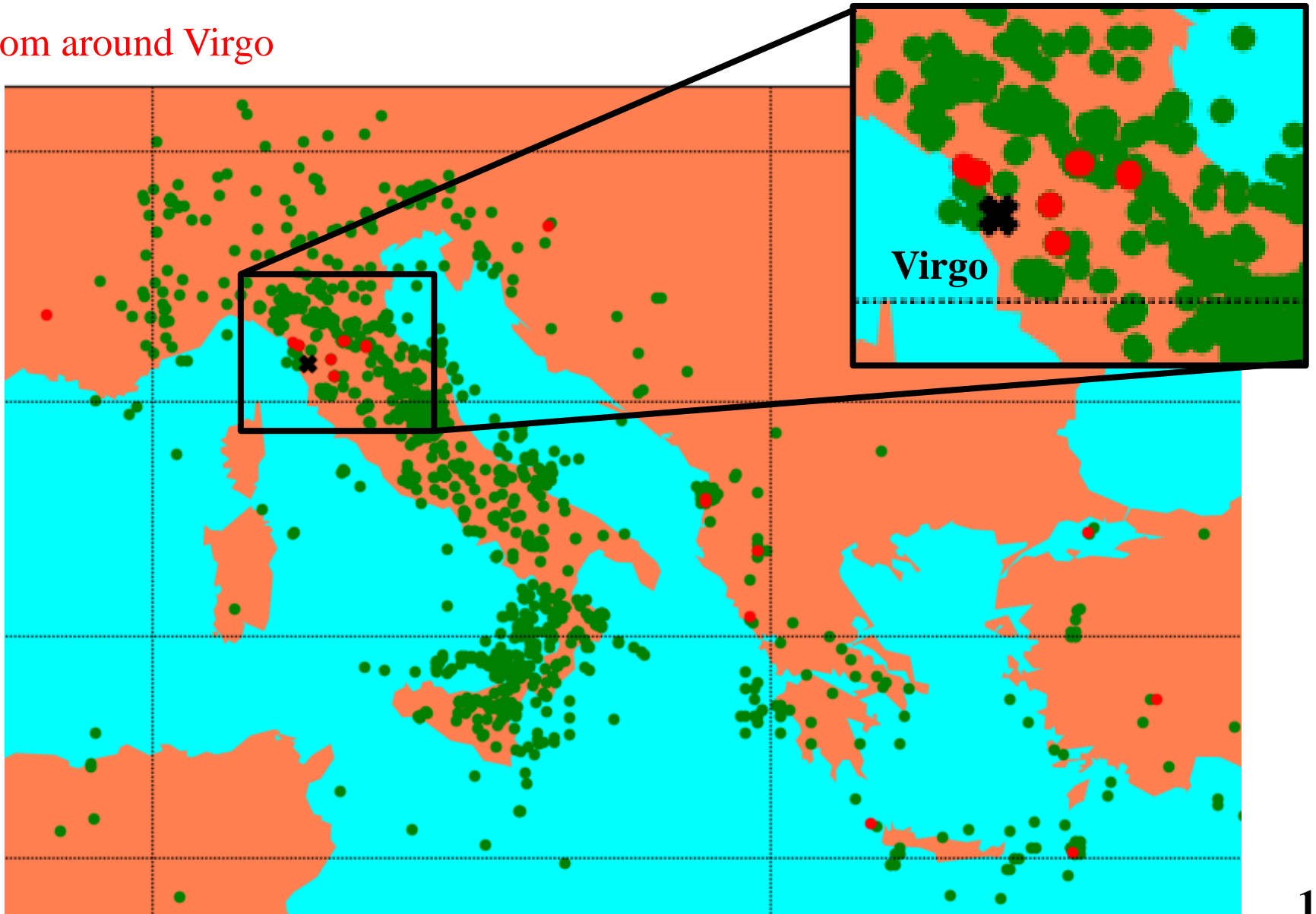
Earthquakes location

- Whole O3
- Excluding earthquakes clearly too weak
 - Empirical cut based on magnitude and distance
- Red dots: lock losses



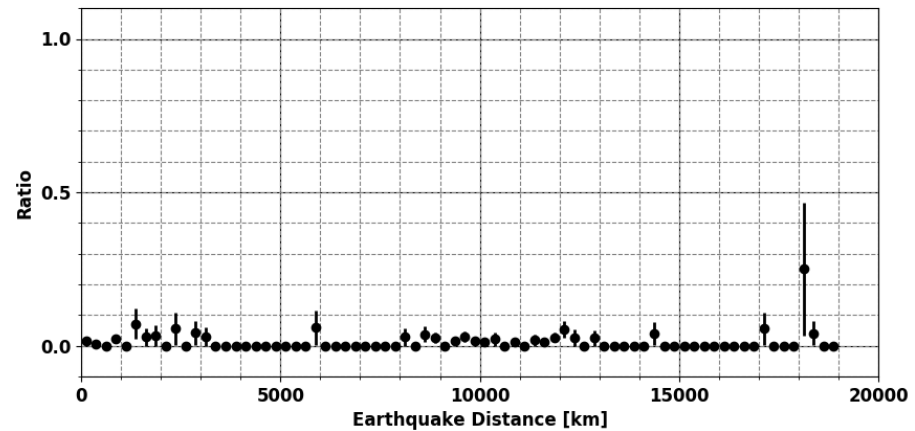
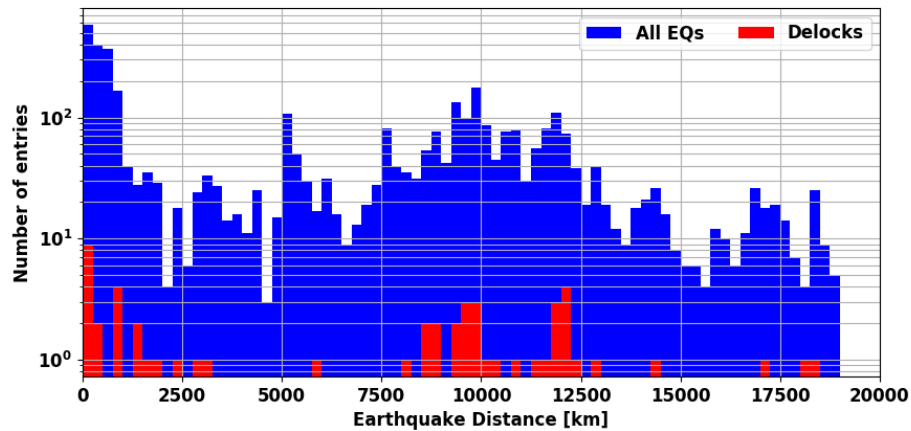
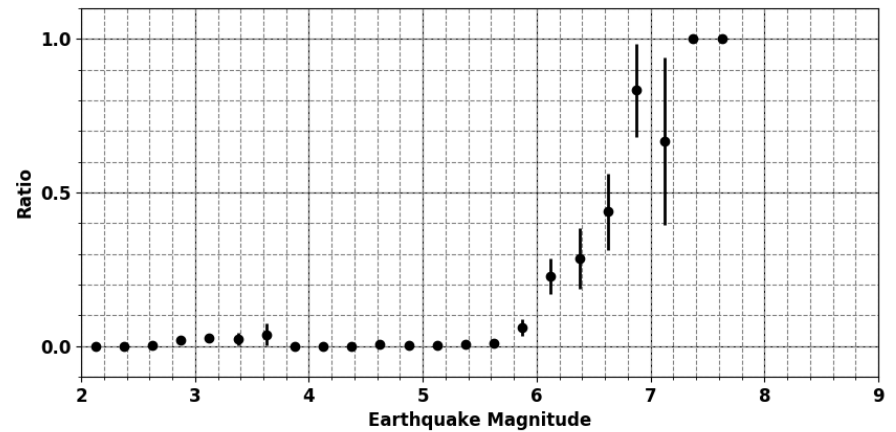
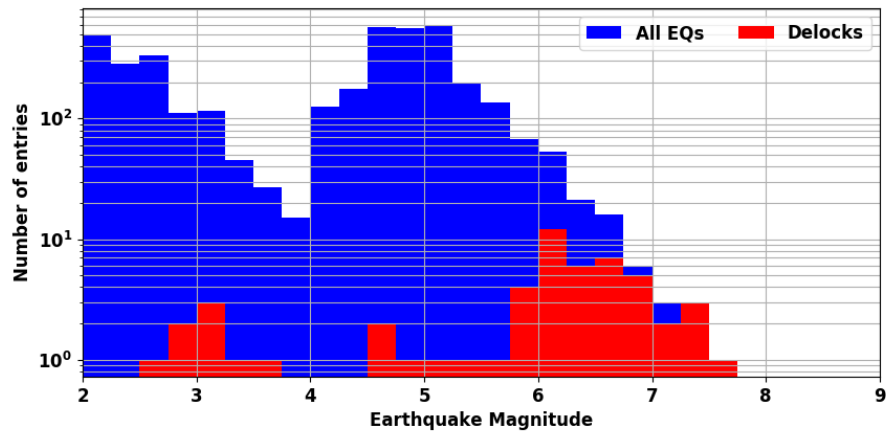
Earthquakes location

- Zoom around Virgo



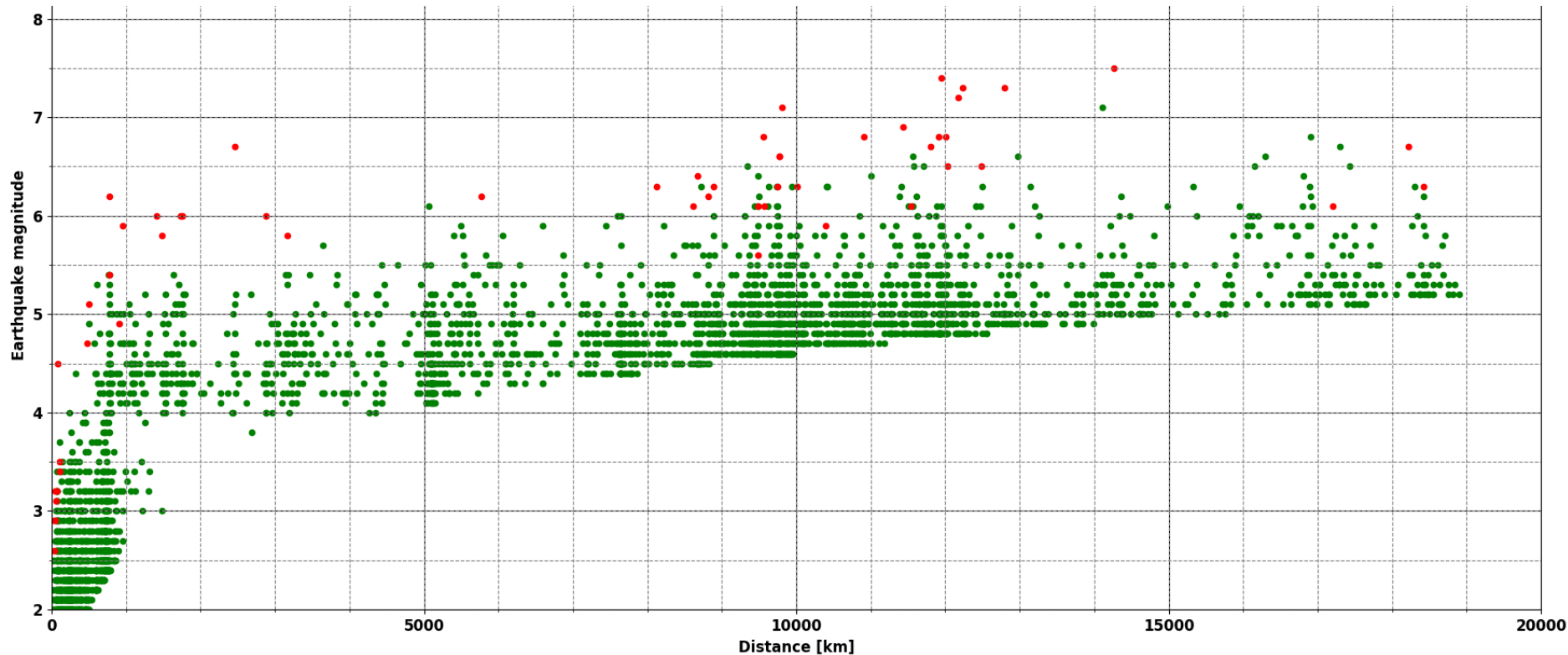
Earthquakes strength

- Classification based on **earthquake magnitude** and **epicenter distance**
 - Blue histograms: all earthquakes
 - Red histograms: earthquakes that led to a control loss during O3



Earthquakes strength

- Classification based on earthquake magnitude and epicenter distance
 - **Green dots**: earthquakes that **did not** led to a control loss
 - **Red dots**: earthquakes that **led** to a control loss

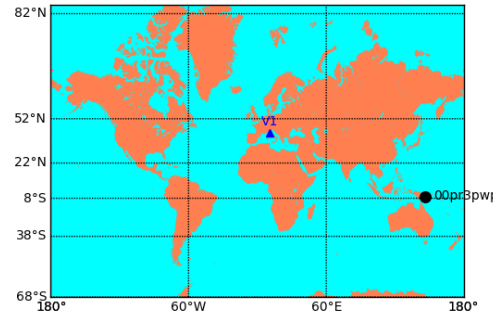


→ **Magnitude** and **distance** are **key parameters**

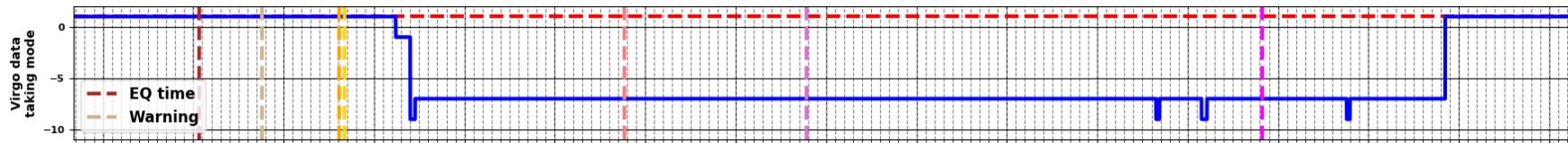
- Others may play a role as well (**epicenter depth**, **azimuth**)
- So probably does the **actual state of the detector** when seismic waves arrive

Riding through a strong earthquake

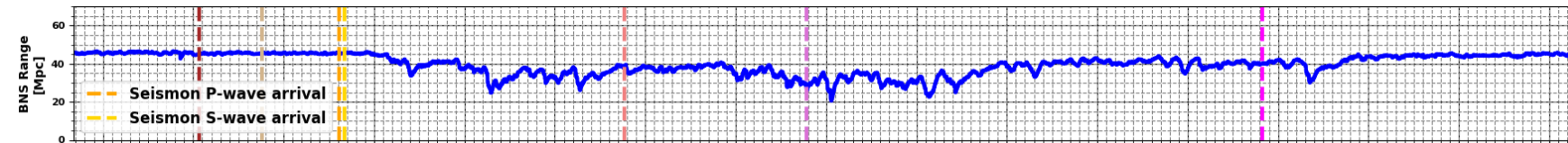
- May 06, 2019
 - Estimated magnitude: 7.1
 - Distance: ~14,000 km



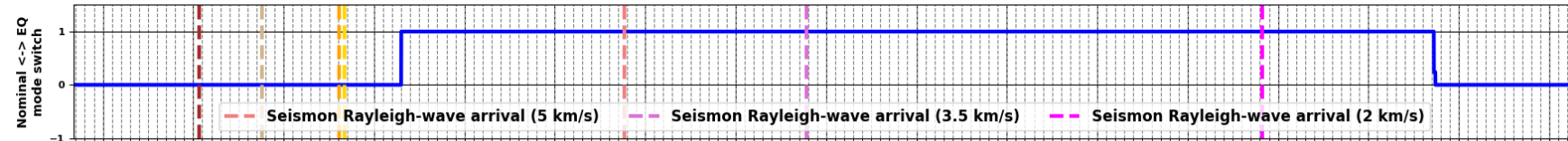
Data Taking Mode



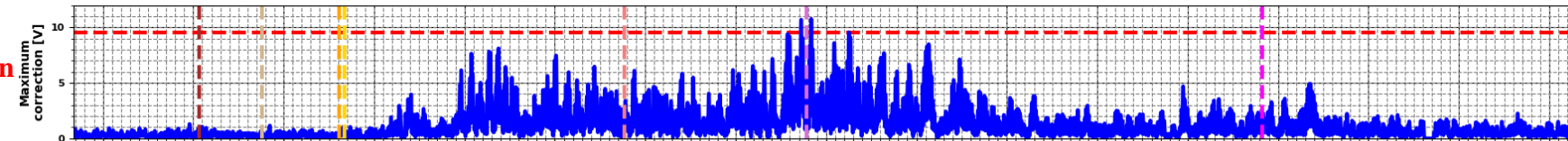
BNS Range



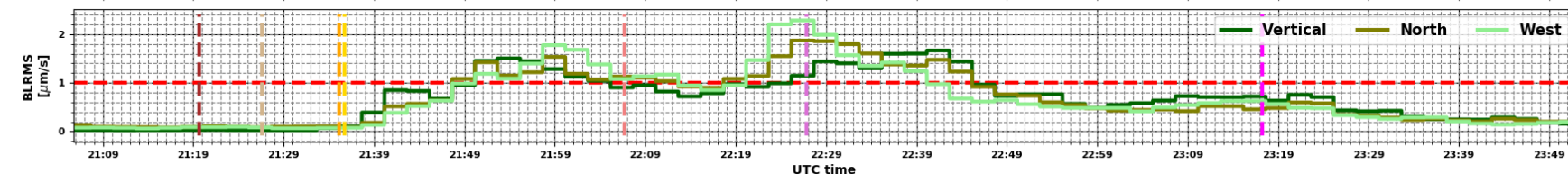
Control Mode



Correction



Seismic noise



→ The only strong earthquake for which switching control mode helped

- By allowing to exceed the correction saturation level w/o losing control

Outlook

- O3 dataset used to quantify findings that had previously been observed qualitatively or empirically
- Seismic noises impact the Virgo detector in a complicated way
 - Various contributions, varying over time for multiple reasons
 - ◆ Weather, human activities
- The Virgo detector appears to be quite robust against those disturbances
 - Limited decrease of BNS range or duty cycle in harsh conditions
- Experience gained to prepare the O4 run
 - More studies needed to understand better the complex behavior of the detector
- Earthquakes are sub-dominant but clear contributors to control losses
 - Two paths explored in parallel for O4
 - Increase the actuation range by dividing the correction force applied among more low-noise actuators
 - Extend the earthquake early warning system
 - ◆ Use the latest Seismon version
 - ◆ Add input stream from INGV (“closer” to Virgo) in addition to USGS