# Advanced Virgo Detector Monitoring and Data Quality

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## Outline

- Advanced Virgo in the GW observatories era
- Current status of Advanced Virgo
- Some Detector Monitoring tools
- The Data Quality of Advanced Virgo
- Gating and Glitch removal

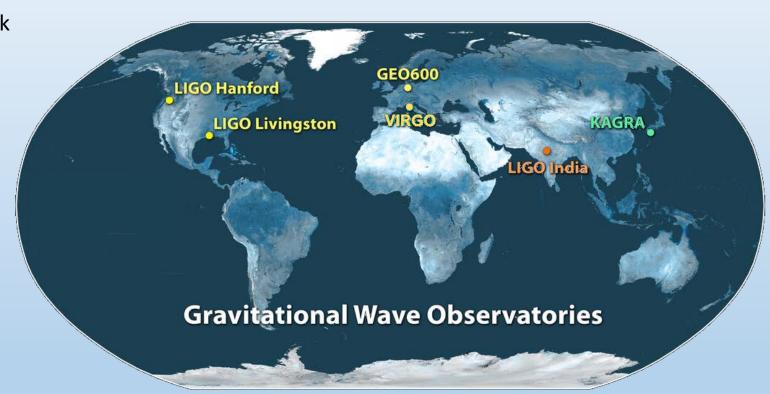
## Advanced Virgo in the GW observatories era

Run O1: first GW detections (GW150914 and GW151226) made with Advanced LIGO detectors and LSC+Virgo analysis

**Advanced Virgo detector** soon included within International GW observatories network for **runs O2 and O3**.

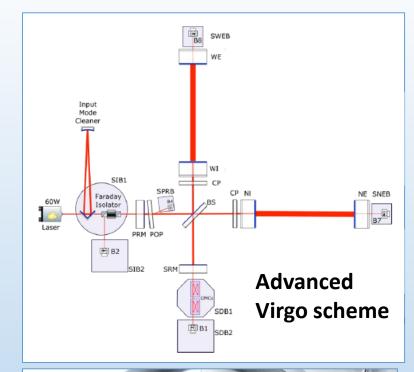
**Detector Monitoring and Data Quality** are important components of Advanced Virgo

to select/confirm the events detected by the **GW observatories network** and to propose them to **EM followup**.



# Advanced Virgo current status

- Installation almost complete
- North arm first lock at the end of April
- Free Michelson data taken mid-June
- Commissioning activity and noise investigation started on subsystems.
- But problems with monolithic suspensions of mirrors.
- Replacing them with steel wires, while investigating.
- Current plan is still to start full interferometer lock and commissioning work before September 2016
- And to be able to participate to run O2...





## **Detector Monitoring and Data Quality**

Detector monitoring and data quality were already quite efficient in previous runs (2007 to 2011).

"The characterization of Virgo data and its impact on gravitational-wave searches", CQG, Vol 29, Nbr 15, 2012

Some components have been improved Some new components are under development.

I will present here only some of those components in which I am more deeply involved.

# Some Detector Monitoring tools

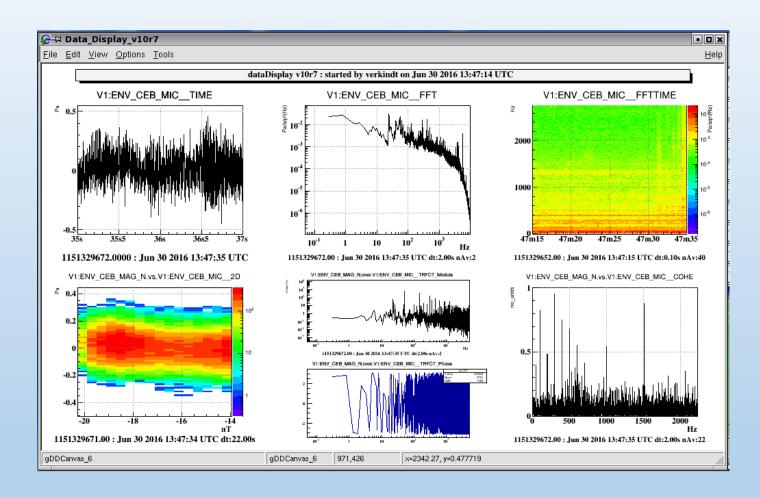
The data Display: to visualize data online

**The MonitoringWeb**: to monitor in-time the trend of various Virgo components

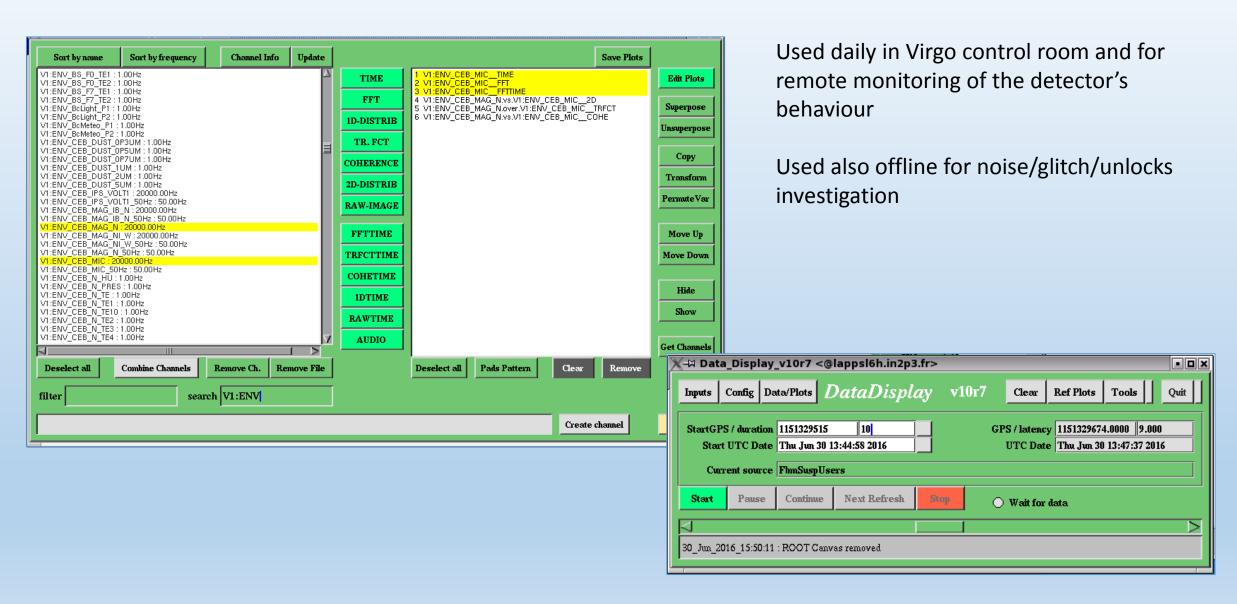
**The Detector Monitoring System**: to monitor online the Virgo detector subsystems

## The data Display

- Software tool based on ROOT
- Read frame formatted data and visualize them offline and online
- Provide various signal processings and various types of plots

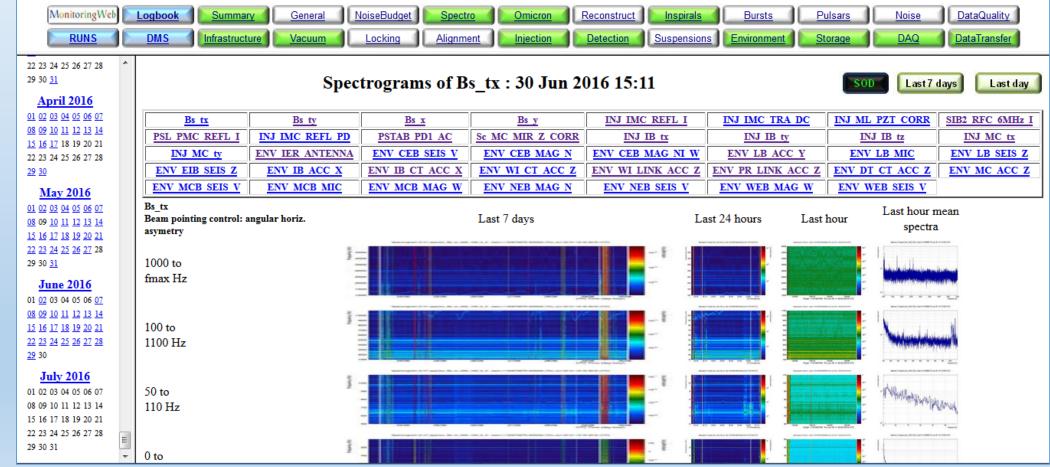


## The data Display



## The MonitoringWeb

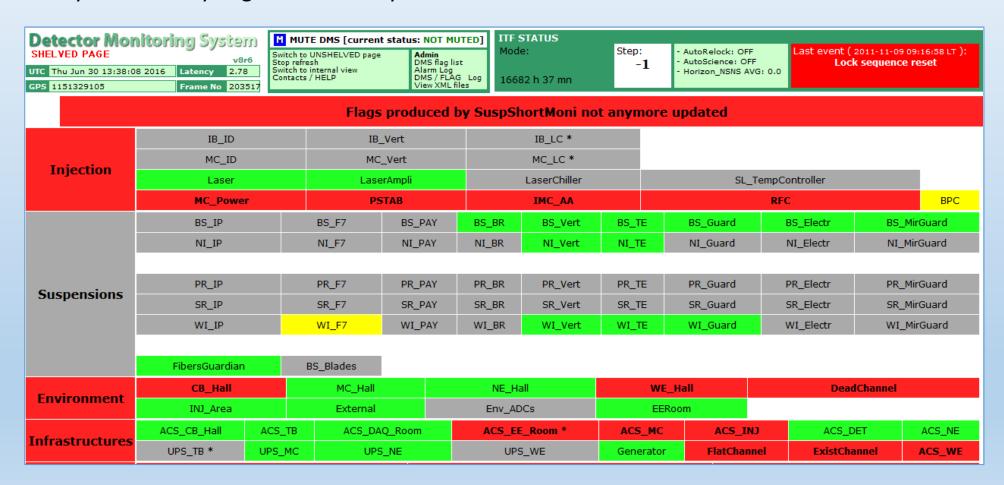
- Similar to the LIGO Summary pages
- Provide online and daily archived plots about the status of various Virgo subsystems
   This includes also spectrograms, data transfer status, data storage status, online data
   quality status, online GW search status, etc...



# The DMS (Detector Monitoring System)

Provides online information with red/green flags and alarms about various parts of the Advanced Virgo detector

Provides also low latency Data Quality flags to be used by the online and offline GW searches



## Advanced Virgo Data Quality

Main complement of detector monitoring is the Data Quality.

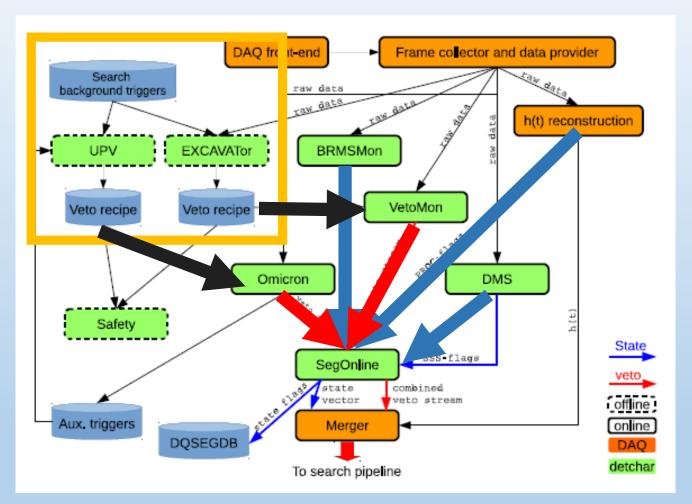
#### Its aims are:

- To provide reliable online quality flags and efficient offline quality flags for O2 and O3 observing runs
- To mitigate any impact of non-stationnary noise coming from sources not removed by commissioning work
- To reduce as much as possible the data analysis backgrounds while keeping low any deadtime not coming from detector unlocks.

## **Advanced Virgo Data Quality**

## Main strategy for online Data Quality:

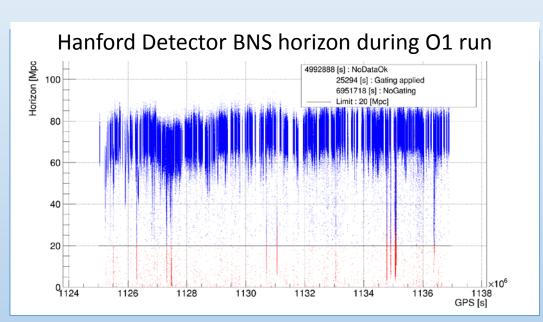
- Provide a generic flag DQ\_VECTOR at 1 Hz
- Provide a set of reliable/efficient analysisdependent vetos at 100 Hz
- Base the online veto production on veto recipes updated periodically.
- Veto recipes set offline using the searches background triggers



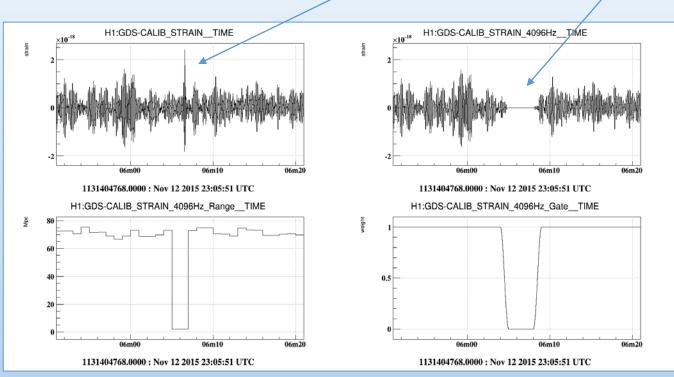
To be added: gating of data and glitch removal (new)

## The Gating of data

Gating: an alternative to the exclusion of data sets by vetos Motivated by the need to keep data flow for the analysis pipelines



**BNS** = Binary Neutron Stars



glitch

AdVirgo Gating developed for MBTA online GW search in O1

One possible step further: **the glitch removal (or mitigation)**Could help to reduce background while preserving as much as possible the percentage of data analyzed

Data put

to zero

C code running a non-linear regression method based on Volterra series and Robust Fast Orthogonal Search

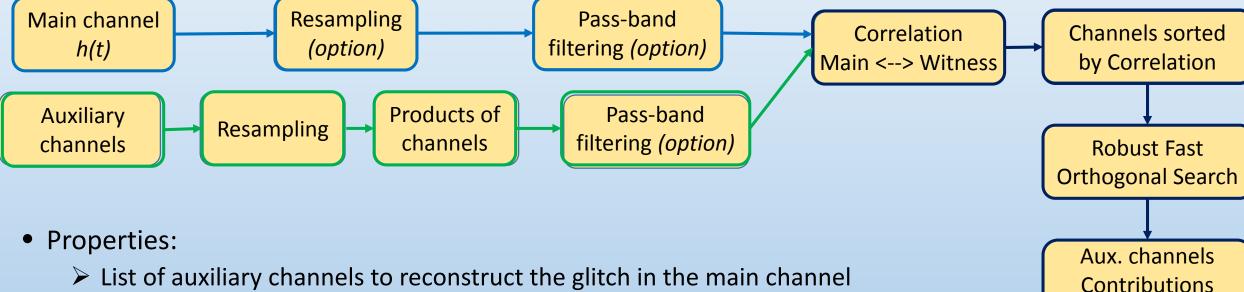
Initially developed in matlab by G. Guidi and F. Piergiovanni Currently developed in C by V. Germain and myself

## Properties:

- > List of auxiliary channels to reconstruct the glitch in the main channel
- ➤ Linear & bi-linear reconstruction → Takes into account the up-conversions
- Orthogonality
- > Estimation of the contribution Q of every auxiliary channel (and product of channels)

## C code running a non-linear regression method based on Volterra series and Robust Fast Orthogonal Search

Initially developed in matlab by G. Guidi and F. Piergiovanni Currently developed in C by V. Germain and myself



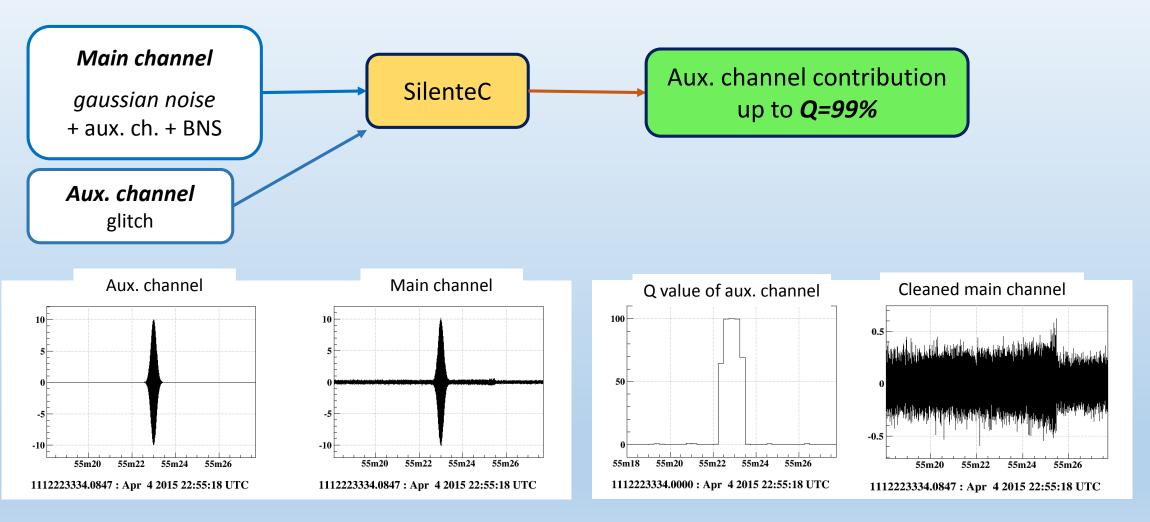
- > Linear & bi-linear reconstruction > Takes into account the up-conversions
- Orthogonality
- > Estimation of the contribution Q of every auxiliary channel (and product of channels)

removed

Clean main

channel created

A simple test: a simulated glitch superposed to a Binary Neutron Star (BNS) signal

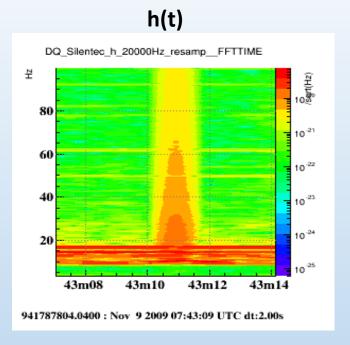


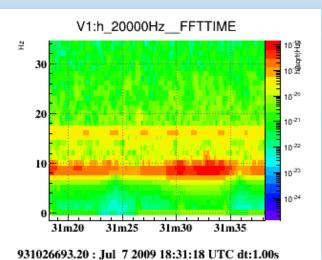
A glitch in VSR2 data associated to a drop of Binary Neutron Star range below 2 Mpc

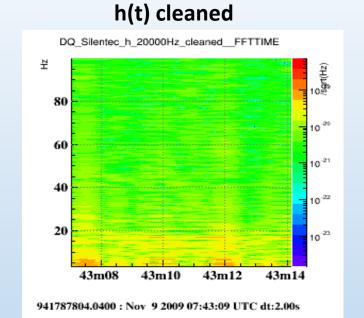
Silentec trying to remove seismic, acoustic and magnetometer channels contributions

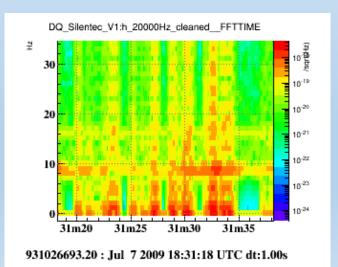
A glitch in VSR2 data associated to an identified seismic noise

Silentec trying to remove all seismic channels contributions









Quite successful cleaning in the frequency range of the glitch

preliminary

Some cleaning in the frequency range of the glitch... but additional noise introduced...

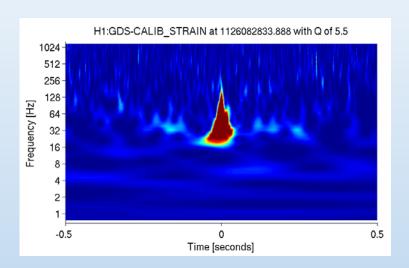
Still investigating

## **Summary**

- Advanced Virgo is coming to full interferometer commissioning phase and is pushing hard to be part of O2 run
- Several monitoring tools are ready or under improvement
- Data Quality strategy is going to be fully implemented and tested
- New tools like « glitch removal » will help noise investigation and will improve data analysis results

# Backup slides

Looking at a **O1 blip glitch**: searching for contributions from Output Mode Cleaner channels and part of Environment Monitoring channels



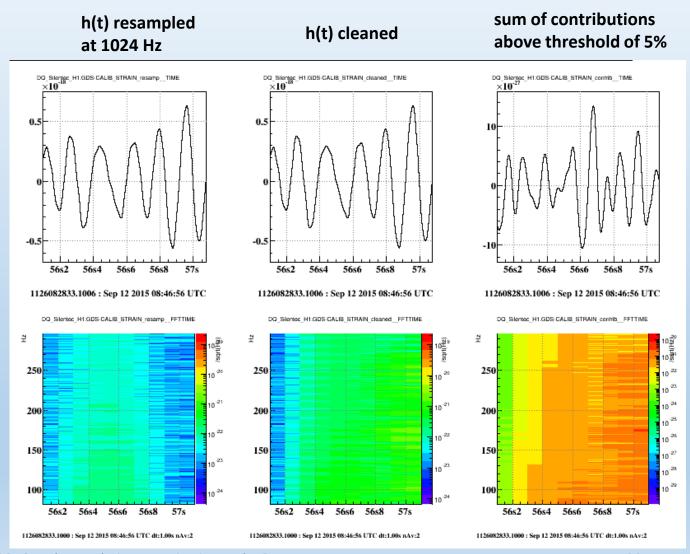
## Threshold at 5%

The main contributing aux. channels are:

[1] H1:OMC-ASC\_Y1\_I\_OUT\_DQ : Q=5.86154%

[2] others : Q=13.7731%

[3] unmodeled : Q=80.3654%



## Virgo Data Quality: Impact on data analysis results

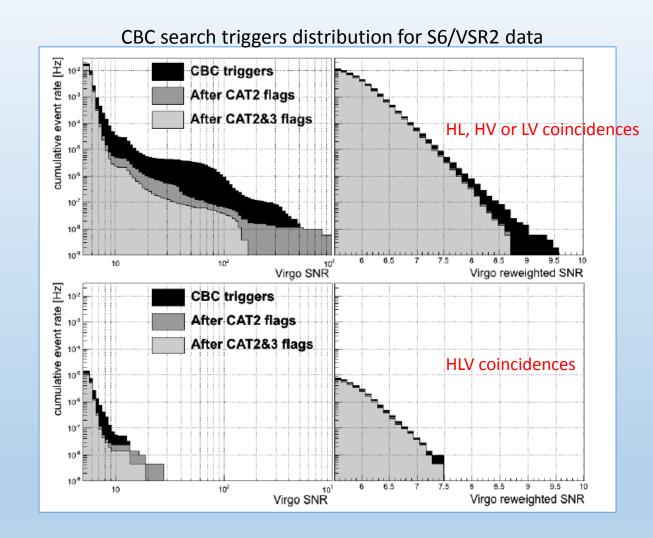
Most of false alarms removed by coincidence with other GW detectors

Remaining coincident events are mainly glitches



Background whose distribution tail lowers the significance of any GW candidate.

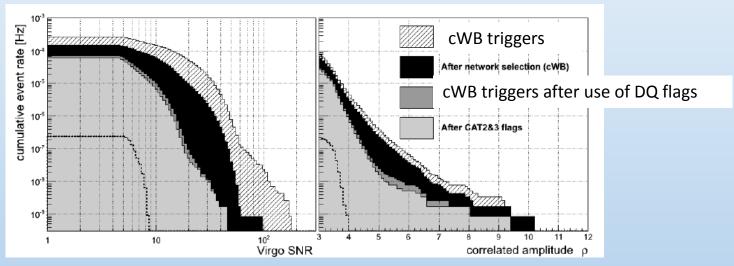
Data Quality vetos reduces this distribution tail.



## **Detector Monitoring and Data Quality**

Detector monitoring and data quality were already quite efficient in previous runs (2007 to 2011).

Effect of DQ flags on the SNR distribution of the cWB burst search triggers for data of the run VSR2 (2009)



From "The characterization of Virgo data and its impact on gravitational-wave searches", CQG, Vol 29, Nbr 15, 2012

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## **Advanced Virgo Data Quality**

### The aims are:

- To provide reliable online quality flags and efficient offline quality flags for O2 and O3 observing runs
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## The main components:

- Data Quality flags based on the Omicron trigger generator and the UPV algorithm
- Data Quality flags based on the Excavator algorithm
- Data Quality flags provided by the Detector Monitoring System
- Alternative or complementary new tools: gating of data and glitch removal