

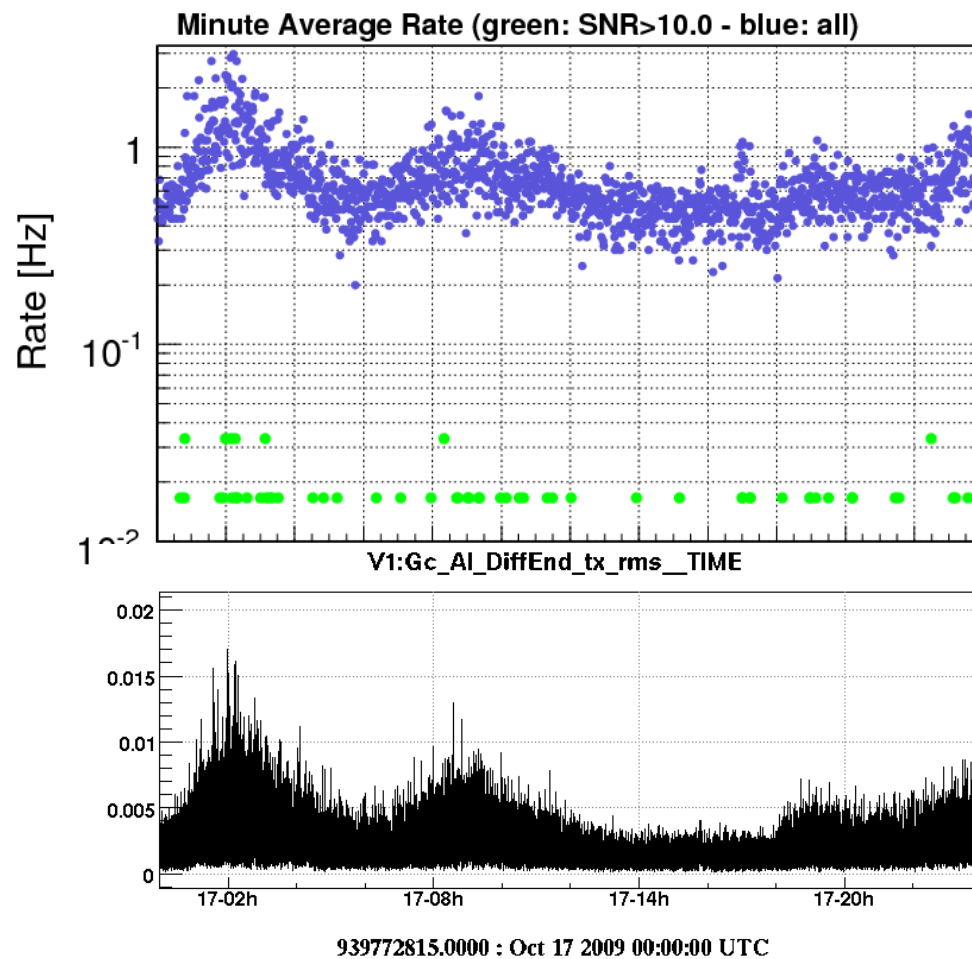
Glitches from alignment fluctuations

&

scattered light

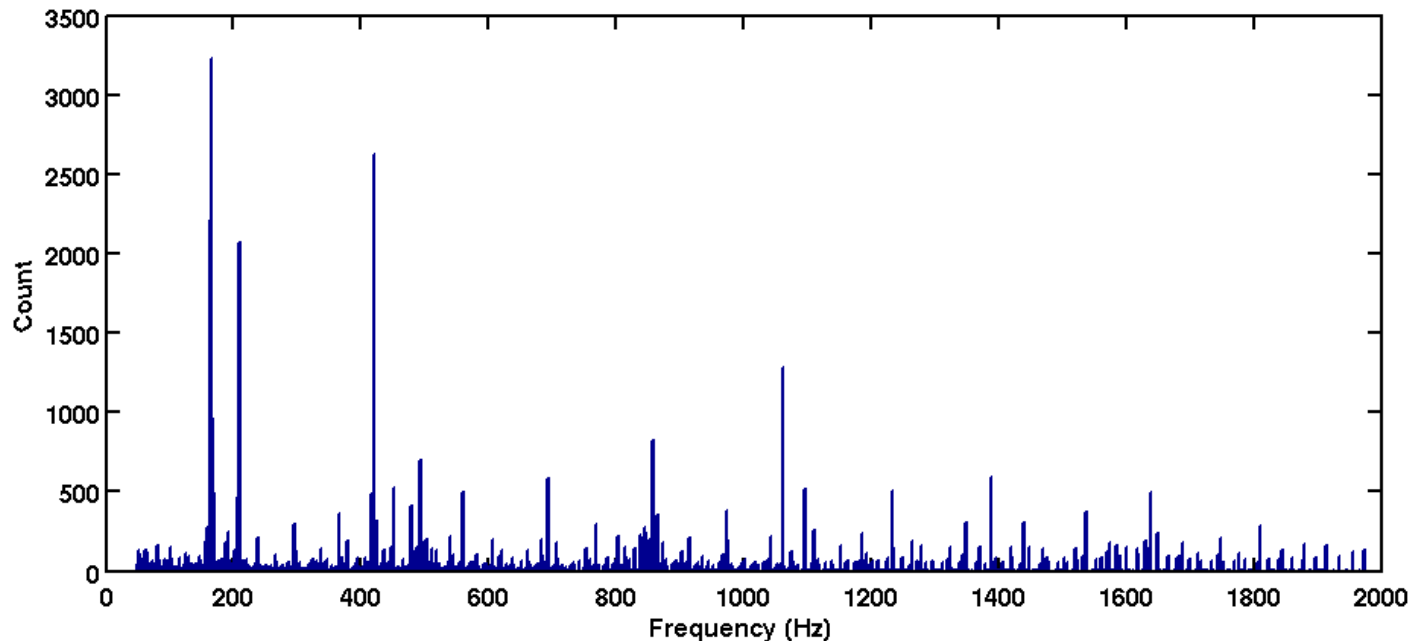
Bas Swinkels

Glitch rate vs DiffEnd_tx



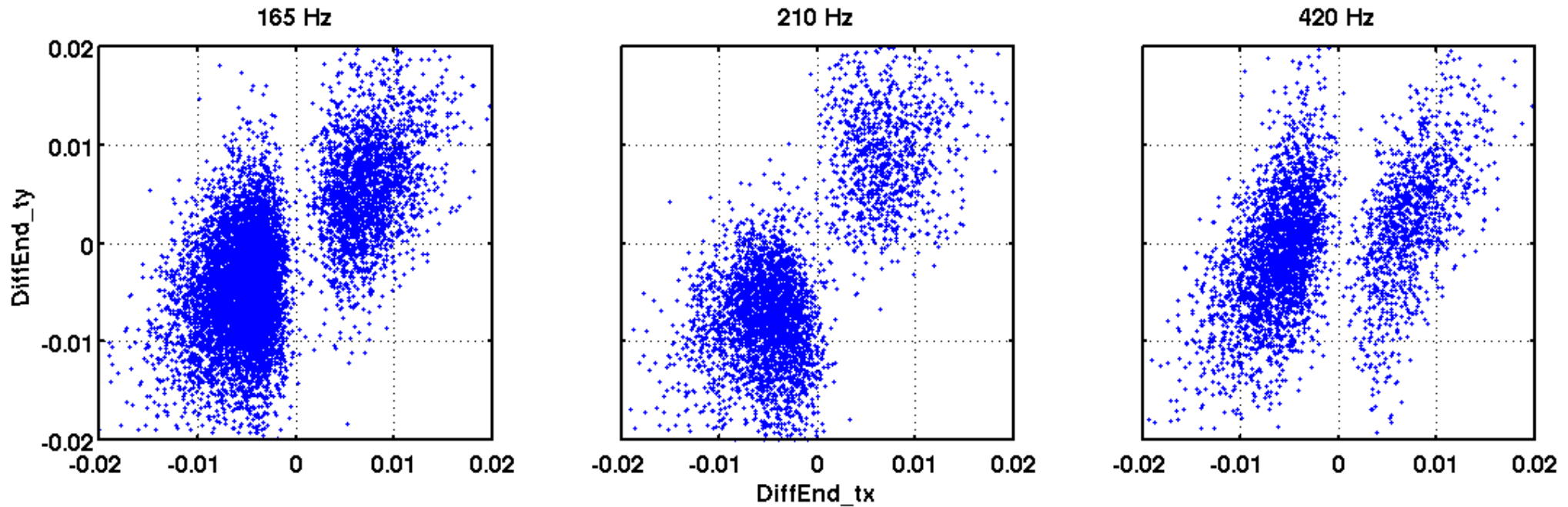
- Glitch rate correlates with accuracy of DiffEnd_tx (Paolo entry #25283)
- Excited by 700 mHz useism, hard to improve controls (#25211)

Glitches vs frequency



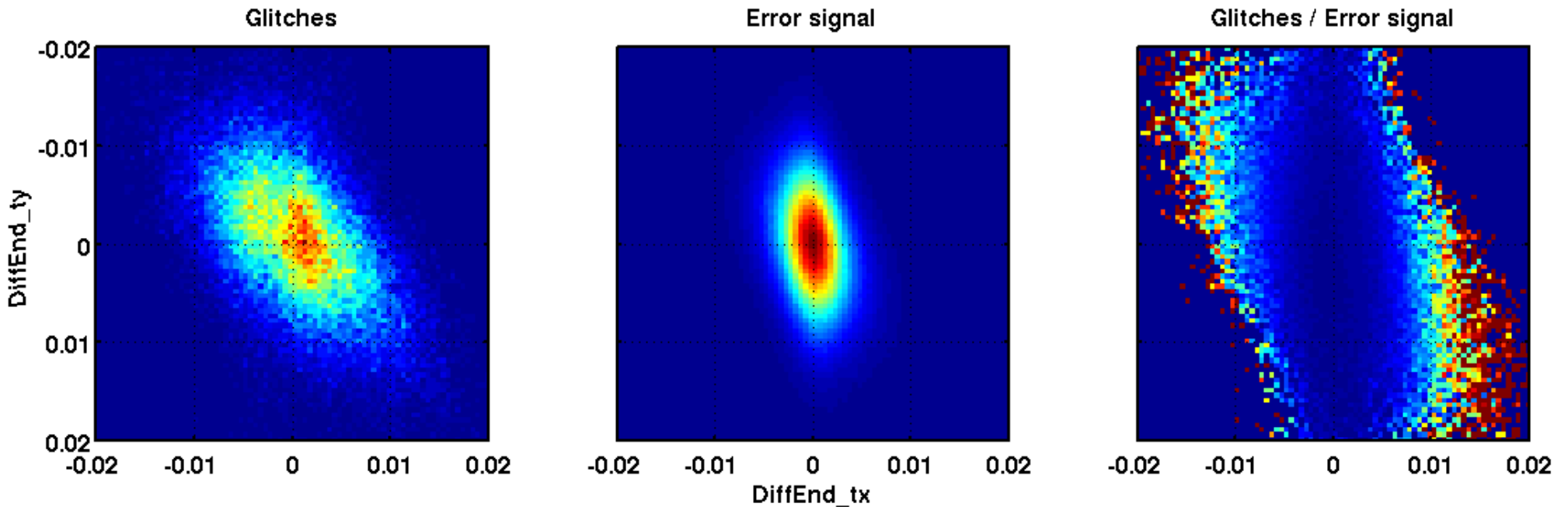
- Histogram of Omega triggers shows many glitches (~20 %) at known resonance frequencies of mounts on suspended detection bench
- 165, 210 and 420 Hz (+some at 800-1000Hz?)
- Mostly low SNR
- Probably not 'real glitches', but noise bumps with non-stationary amplitude

Glitches vs DiffEnd position



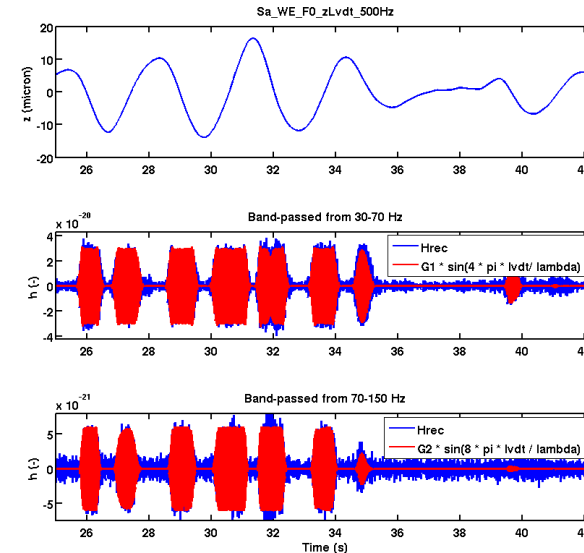
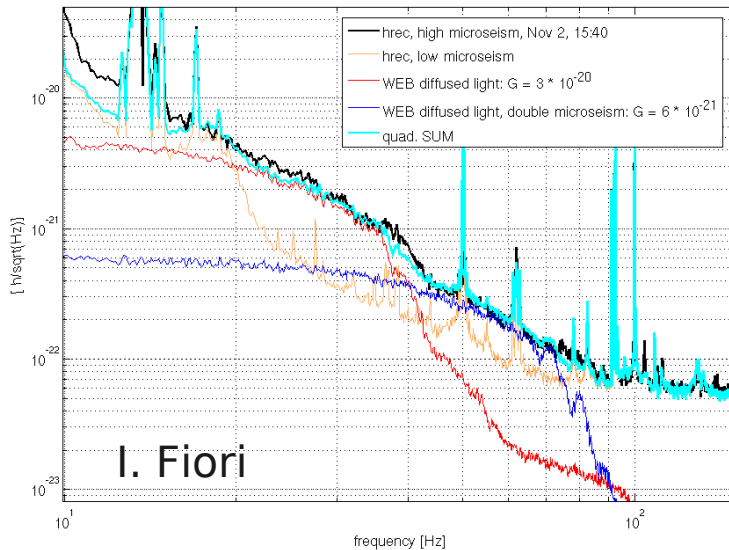
- Alignment position at moment of glitch shows different patterns for different frequencies
- DiffEnd alignment controls 'darkness' and size of B1
- Each frequency might correspond to clipping/scattering by different amount
- Easy solution might be to throw away all glitches at these frequencies, but DQ people have moral problems with that

2D histograms



- Patterns not visible when plotting glitches at all frequencies, but cloud of glitches larger than expected from error signal
- Glitch probability depends mainly on tx
- Quick test using threshold on $\text{abs}(\text{DiffEnd_tx})$ shows 12% efficiency for 1% of dead-time (threshold = 0.008)
- More tests needed

Scattered light



- Well understood problem, see old presentations by Fiori / Tournefier
- Biggest problem at WE, due to higher mirror transmission
- Single bounce

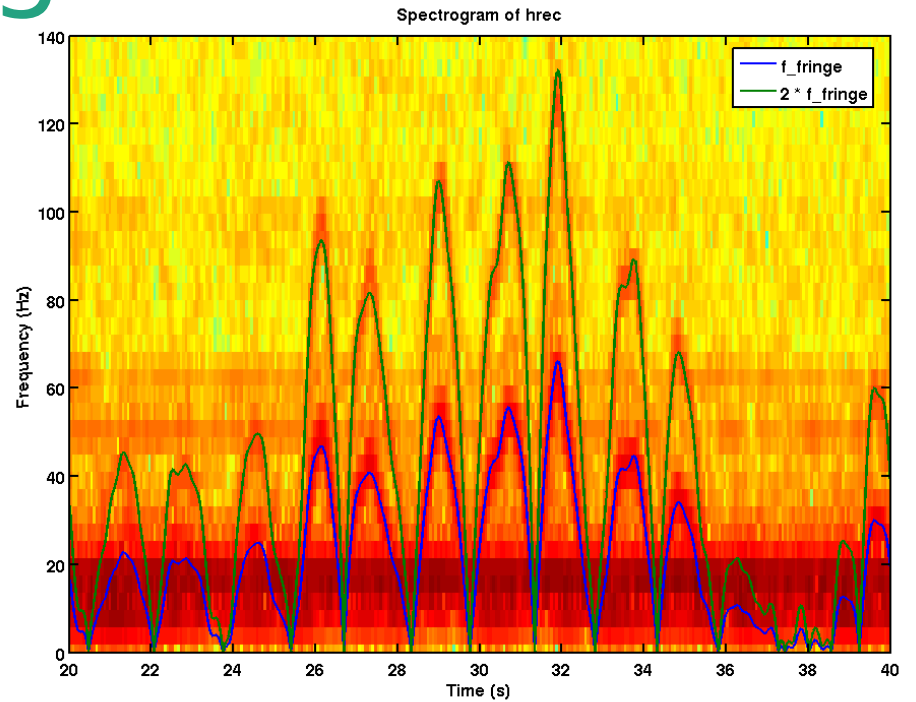
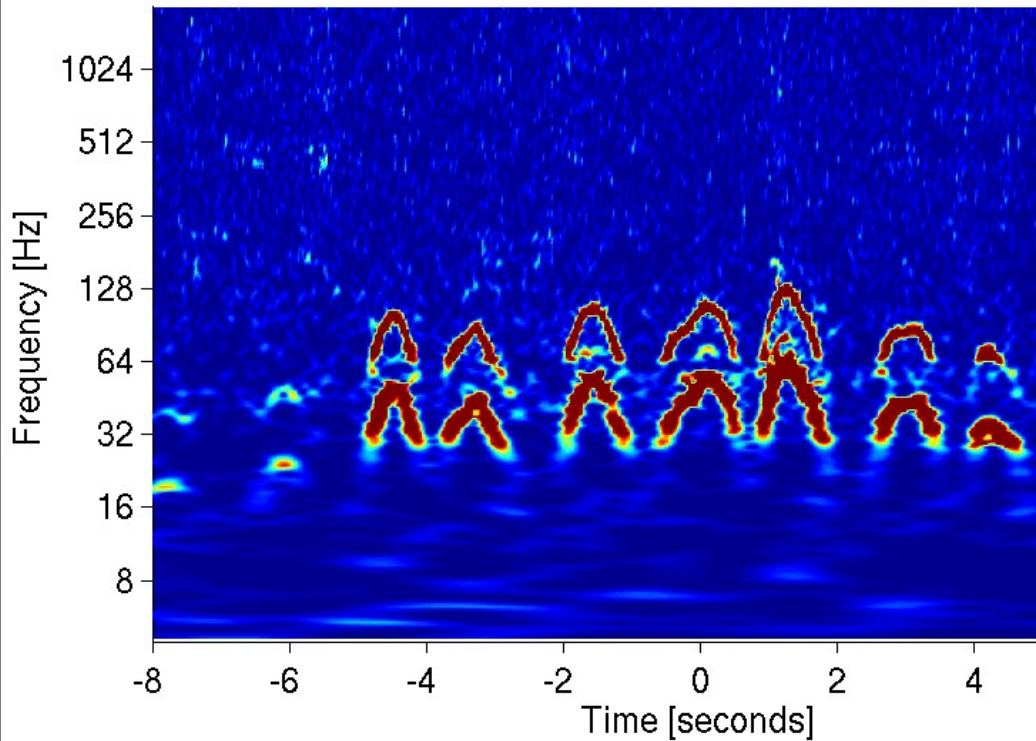
$$h(t) = G1 * \sin(4 * \pi * x(t) / \lambda)$$

- Double bounce

$$h(t) = G2 * \sin(8 * \pi * x(t) / \lambda)$$

- Model works well in both time and frequency domain

Spectrograms



- Arched fringes in spectrogram, similar to swinging baffles
- 1 cycle in hrec when scatter path changes by 1 fringe, so fringe frequency proportional to velocity

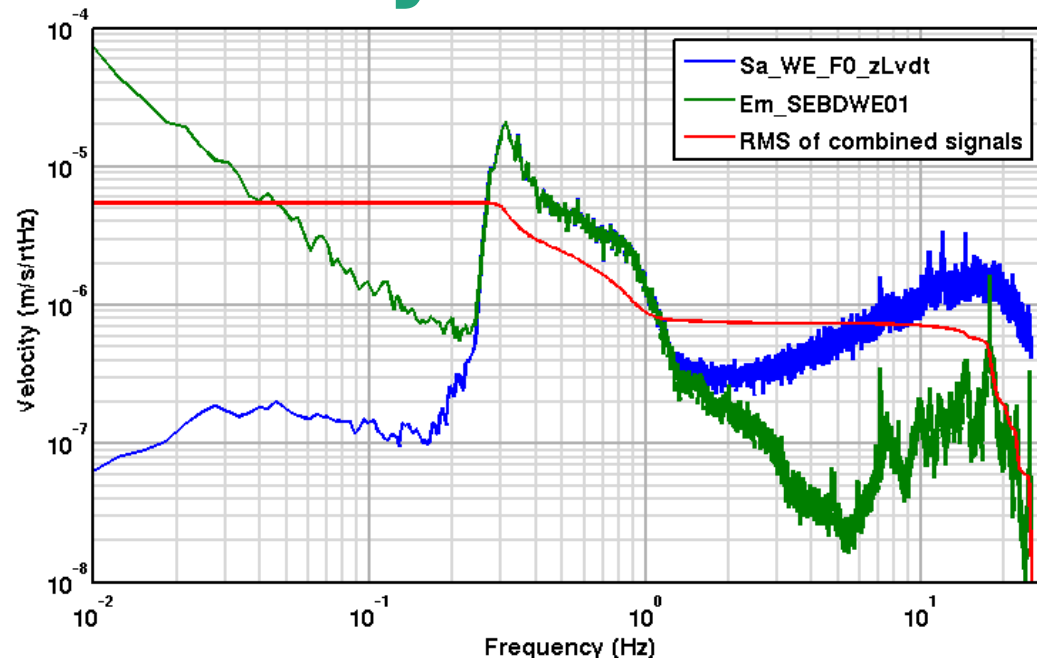
$$f_{\text{fringe}} = 2 * \text{abs}(V / \text{lambda})$$

- 'Double bounce' path

$$f_{\text{fringe}} = 4 * \text{abs}(V / \text{lambda})$$

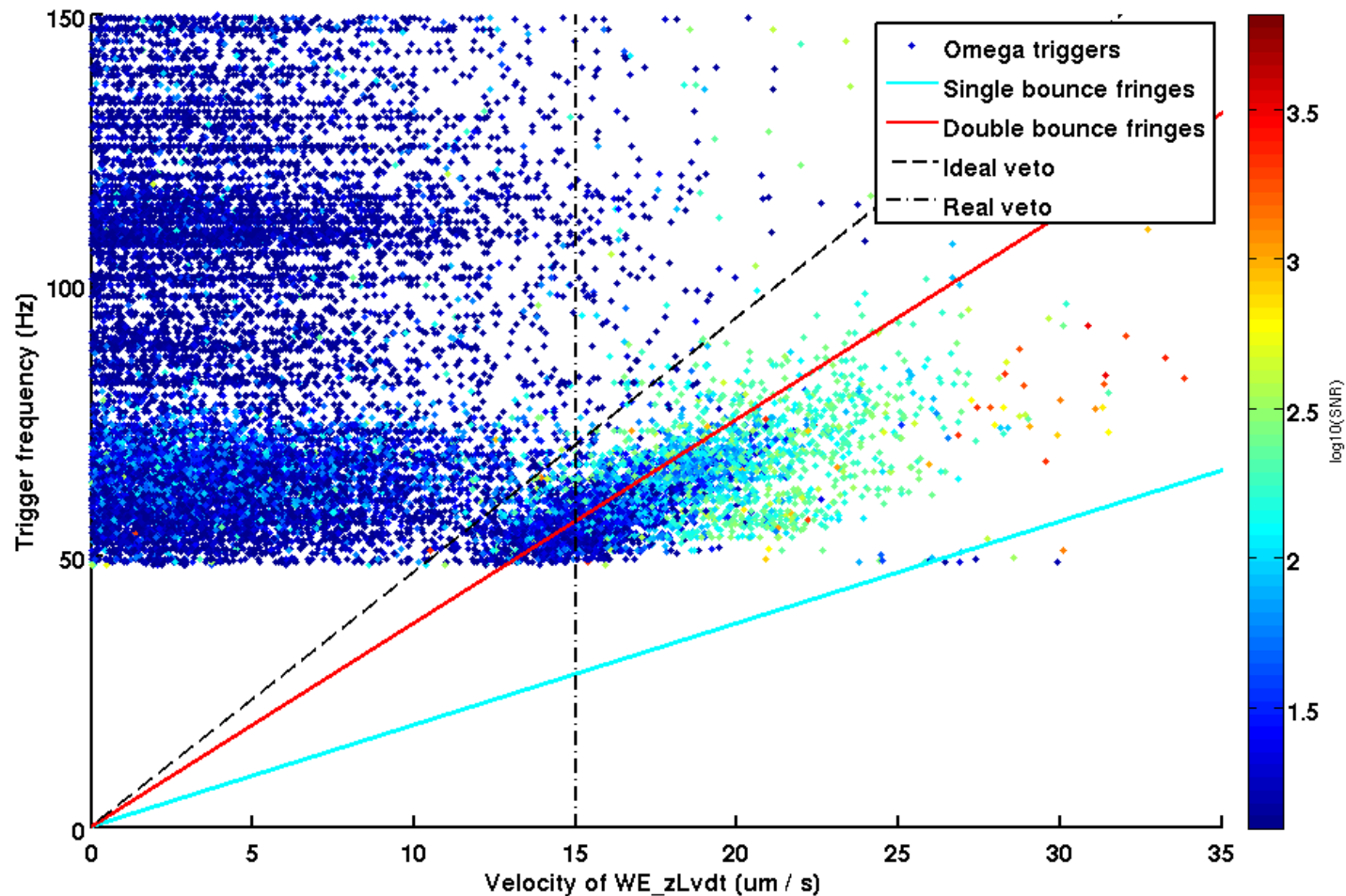


Velocity estimation



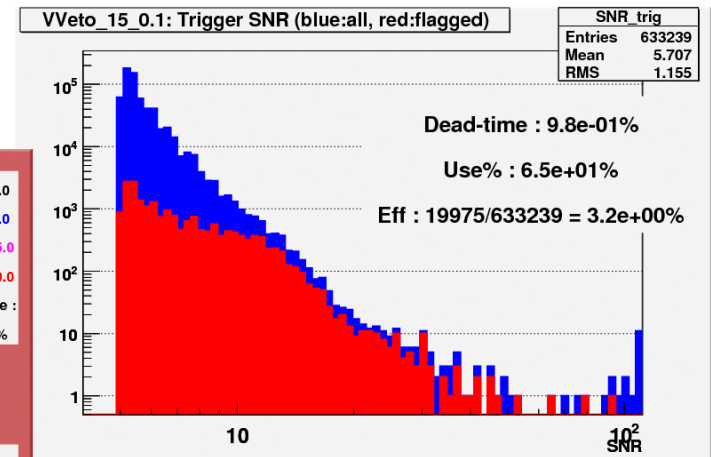
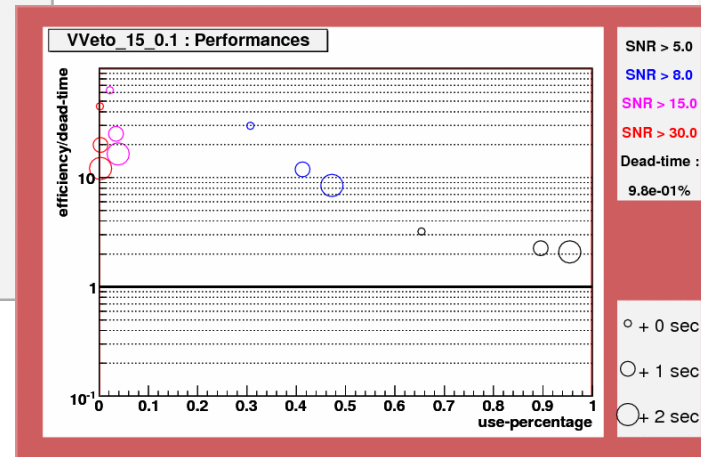
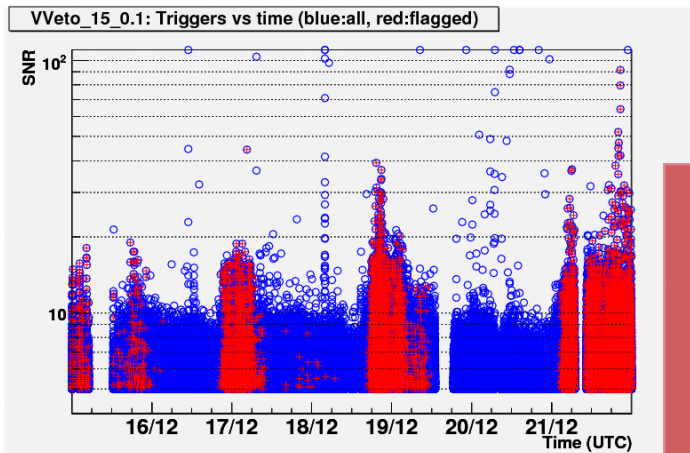
- End-mirror is fixed in space by controls, rest of the world moving at useism frequencies (0.25 - 1Hz) when weather is bad
- Seen by several sensors
 - IP top stage position sensor: Sa_WE_F0_zLvd
 - Local control camera: Gx_WE_z
 - Bench seismometer: Em_SEBDWE01
- LVDT works best for determining velocity

Glitch frequency vs velocity



- Distinct cloud of points due to 'double bounce fringes'
- Ideal veto based on frequency and velocity, but 'unsafe'
- Real veto based on velocity only

Veto performance



- New sub-second veto based on velocity implemented by F. Robinet http://wwwcascina.virgo.infn.it/DataAnalysis/Burst/dqtest/perf_SE/
- Very selective, high efficiency/dead-time
- Might be used as CAT 2 veto
- Kills most high SNR glitches, but does not solve all scattered light and useism problems
- Old veto based on BRMS (0.25-1Hz) of seismometers will probably remain as CAT 3 veto to flag periods with high useism

Concluding

- Family of glitches at frequency of output bench resonances, modulated by DiffEnd alignment fluctuations
 - possible veto to be tested
- WE scattered light glitches
 - glitch frequency well modeled by velocity
 - veto based on velocity only shows good performance
 - might efficiently kill a class of big glitches, but no silver bullet for scattered light problems
- Also look at $> 1\text{ Hz}$ value of signals instead of $\leq 1\text{ Hz}$ BRMS only!
- In case of known problems: make histograms vs relevant parameters
- Check other DOFs, nothing found so far

End

Backup

