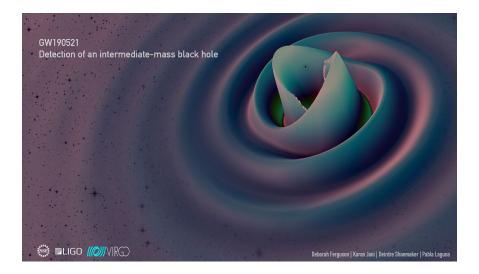
# GW190521: A Binary Black Hole Merger with a Total Mass of 150 Solar Masses

LIGO Scientific Collaboration and Virgo Collaboration



Discovery paper -Phys. Rev. Lett. 125, 101102 (2020) https://doi.org/10.1103/PhysRevLett.125.101102

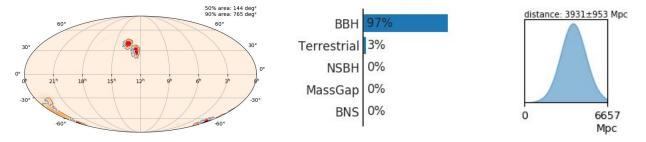
(Astro)physical implications -Astroph. J. Lett 900 L13 (2020)

https://doi.org/10.3847/2041-8213/aba493

Data release https://dcc.ligo.org/LIGO-P2000158/public

# **Discovery of GW190521**

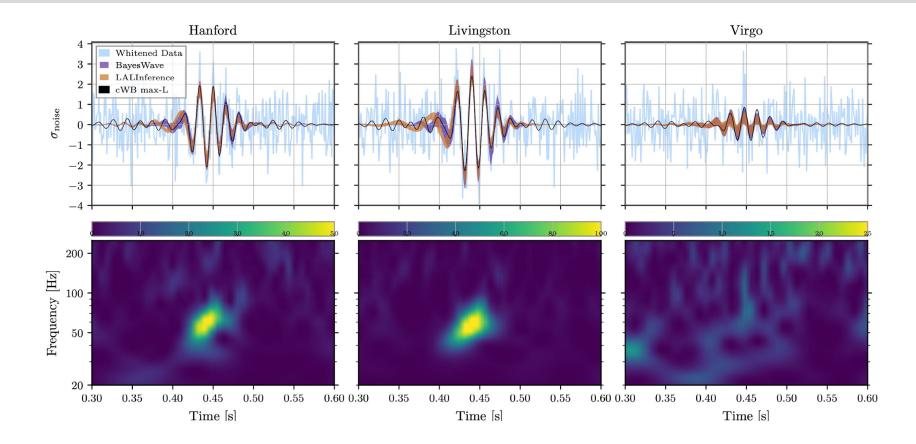
Event was detected May 21, 2019 at 03:02:29 UTC and publicly reported (as S190521g) 6 minutes later.



Subsequent offline analysis confirmed that this is a confident gravitational wave detection by LIGO and Virgo.

Parameter estimation routines using more complex general relativistic models for the waveform provide the information about the signal source.

# GW190521 in LIGO Hanford, LIGO Livingston, Virgo



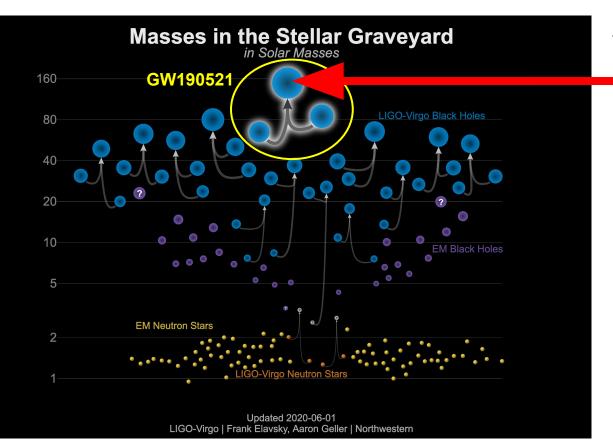
# **GW190521 Parameters**

- Most massive observation to date
- Most distant
- Pair-instability supernova mass gap, 65-120 M₀
- Intermediate Mass Black Hole
- Important astrophysical implications
- Orbital precession

TABLE I. Parameters of GW190521 according to the NRSur7dq4 waveform model. We quote median values with 90% credible intervals that include statistical errors.

Parameter	
Primary mass	$85^{+21}_{-14} M_{\odot}$
Secondary mass	$66^{+17}_{-18} M_{\odot}$
Primary spin magnitude	$0.69^{+0.27}_{-0.62}$
Secondary spin magnitude	$0.73^{+0.24}_{-0.64}$
Total mass	$150^{+29}_{-17} M_{\odot}$
Mass ratio $(m_2/m_1 \le 1)$	$0.79^{+0.19}_{-0.29}$
Effective inspiral spin parameter $(\chi_{eff})$	$0.08\substack{+0.27\\-0.36}$
Effective precession spin parameter $(\chi_p)$	$0.68^{+0.25}_{-0.37}$
Luminosity Distance	5.3 <sup>+2.4</sup> <sub>-2.6</sub> Gpc
Redshift	$0.82^{+0.28}_{-0.34}$
Final mass	$142^{+28}_{-16} M_{\odot}$
Final spin	$0.72^{+0.09}_{-0.12}$
$P (m_1 < 65 M_{\odot})$	0.32%
log <sub>10</sub> Bayes factor for orbital precession	$1.06\substack{+0.06\\-0.06}$
log <sub>10</sub> Bayes factor for nonzero spins	$0.92^{+0.06}_{-0.06}$
log <sub>10</sub> Bayes factor for higher harmonics	$-0.38^{+0.06}_{-0.06}$

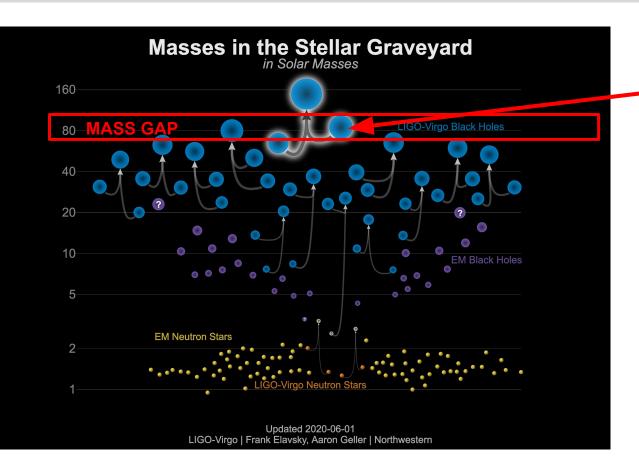
### The most massive black hole ever observed with gravitational waves



The final black hole is

- the most massive black hole ever observed with gravitational waves
- the first evidence of a black hole in the 100 -1000 solar mass range
- an intermediate-mass black hole: the missing link between stellar-mass and supermassive black holes

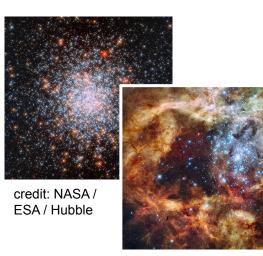
## The first black hole in the pair-instability mass gap



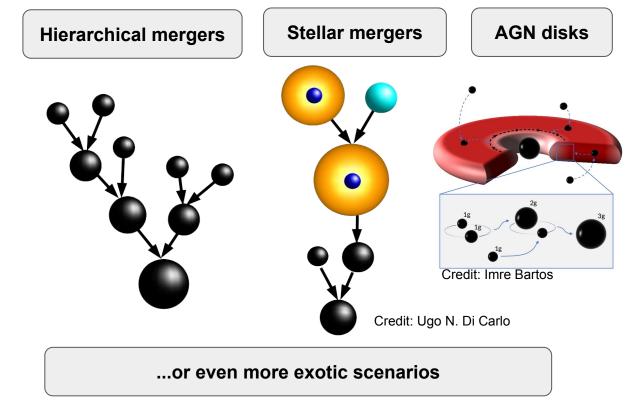
- One of the two merging black holes has mass 85 solar masses: it cannot form from stellar collapse
- Very massive stars (He core ~ 30 - 135 solar masses) undergo (PULSATIONAL) PAIR INSTABILITY
- Expected gap in the black hole mass spectrum between ~ 65 and ~120 solar masses

# Challenge for the models of black hole formation

In dense star clusters and galactic nuclei, black holes can have close encounters with other black holes



credit: NASA, ESA, F. Paresce, R. O'Connell



# Decoding the black holes

#### State-of-the-art tools:

 3 advanced GR models + 3000 supercomputer simulations of black hole collisions from Einstein's GR theory

#### **Records book event!**

- Heaviest
- First of its kind
- Most energetic
- Farthest (halfway since the Big Bang)
- Tilted spins (hints of the origins story)

# New Tests for Einstein

#### This event is a rare catch:

- Only once every 10 years in Gpc<sup>3</sup>
- 500 times rarer than past LIGO-Virgo events

#### Full points to Einstein!

- Consistency at extreme gravity regime between pre- and post-merger black holes
- No signs in the data for alternative theories

pre-merger

post-merger

# Future is bright!

### Black hole desert:

- LIGO-Virgo-KAGRA will have large sample of GW190521-like black holes
- Next-generation detectors can find all size of intermediate-mass black holes

### Multi-band source:

- Space+ground
- Critical for origin stories

