

OMC and dark fringe characterisation

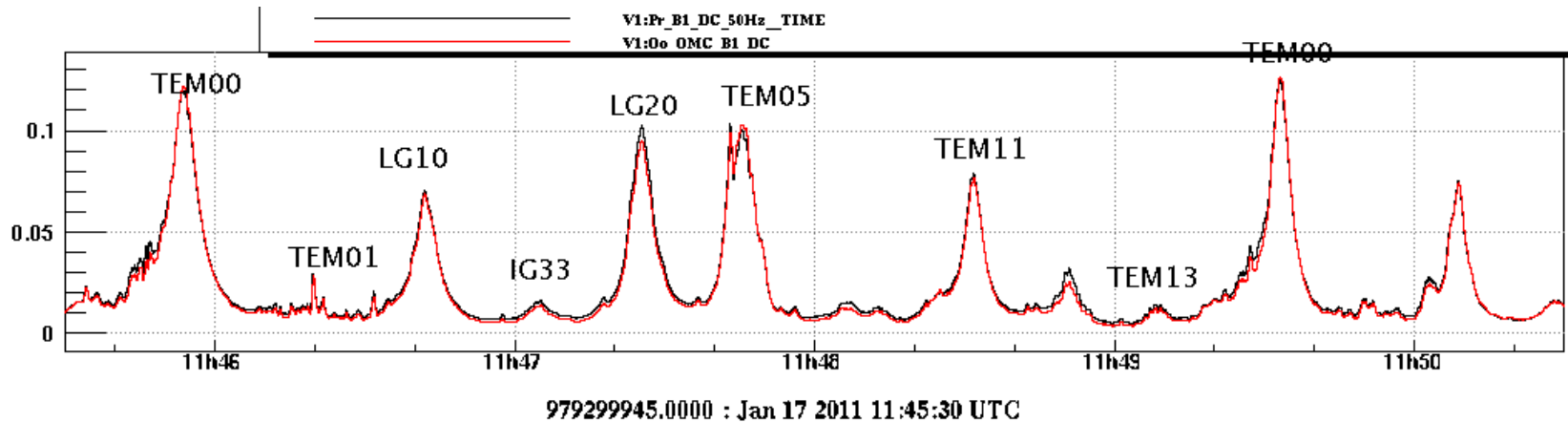
Edwige Tournefier, Raffaele Flaminio, Romain Gouaty

Virgo week - Commissioning meeting
Feb 7th, 2011

- OMC scans
- OMC matching and ITF alignment
- Dark fringe composition: higher order modes, carrier/sidebands
- OMC and Fabry-Perot higher order modes transmission
- Measurement of FP losses

OMC scans: what for?

- Example of OMC scan (with CHRoCC ~310 degrees, logbook entry 28635 by Julien)



These data are used to extract the dark fringe content in HO modes and to measure the OMC matching:

- ⇒ Which part of the observed HOMs is due to OMC mismatching / misalignment?
- ⇒ Which part is really the dark fringe content?

Look at the GW signal in each HO mode:

- modes due to OMC mismatching/ misalignment contain GW signal
- dark fringe HO modes do not contain GW signal

OMC scans in practice

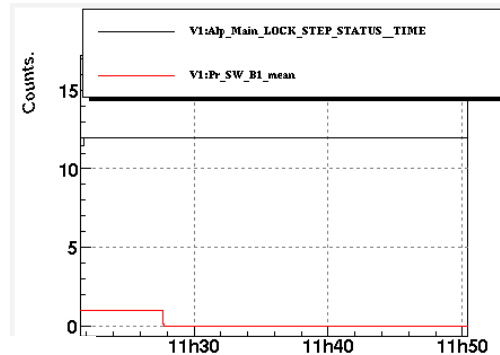
- ITF in step 8 or 12
- Dark fringe controlled with B1p
 - Large offset on B1p_ACp due to the presence of HO mode
 - ⇒ The differential working point is not the good one
 - ⇒ B1_ACp often saturates when the ITF is controlled with B1p

This offset should be subtracted

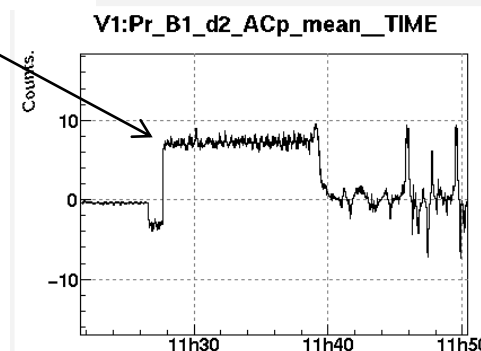
⇒ Easier interpretation of OMC scans

⇒ Easier switch B1p<->B1

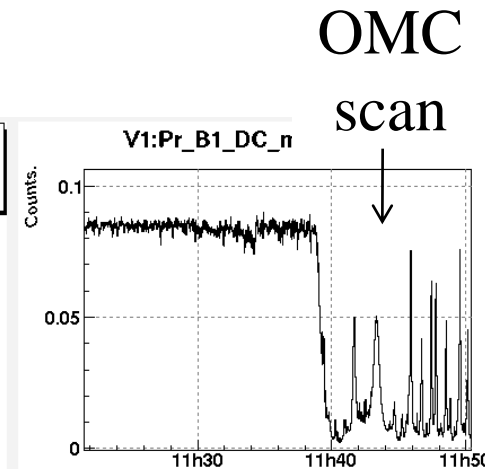
DARM control on B1p



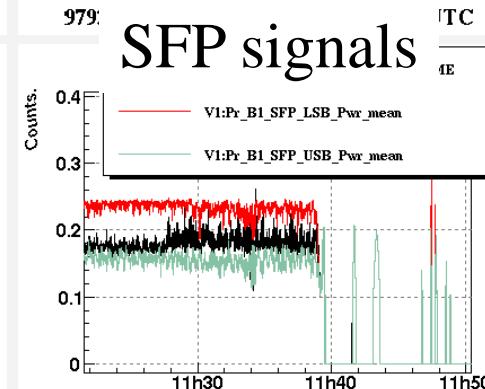
298505.0000 : Jan 17 2011 11:21:30 UTC



979298505.0000 : Jan 17 2011 11:21:30 UTC



979:



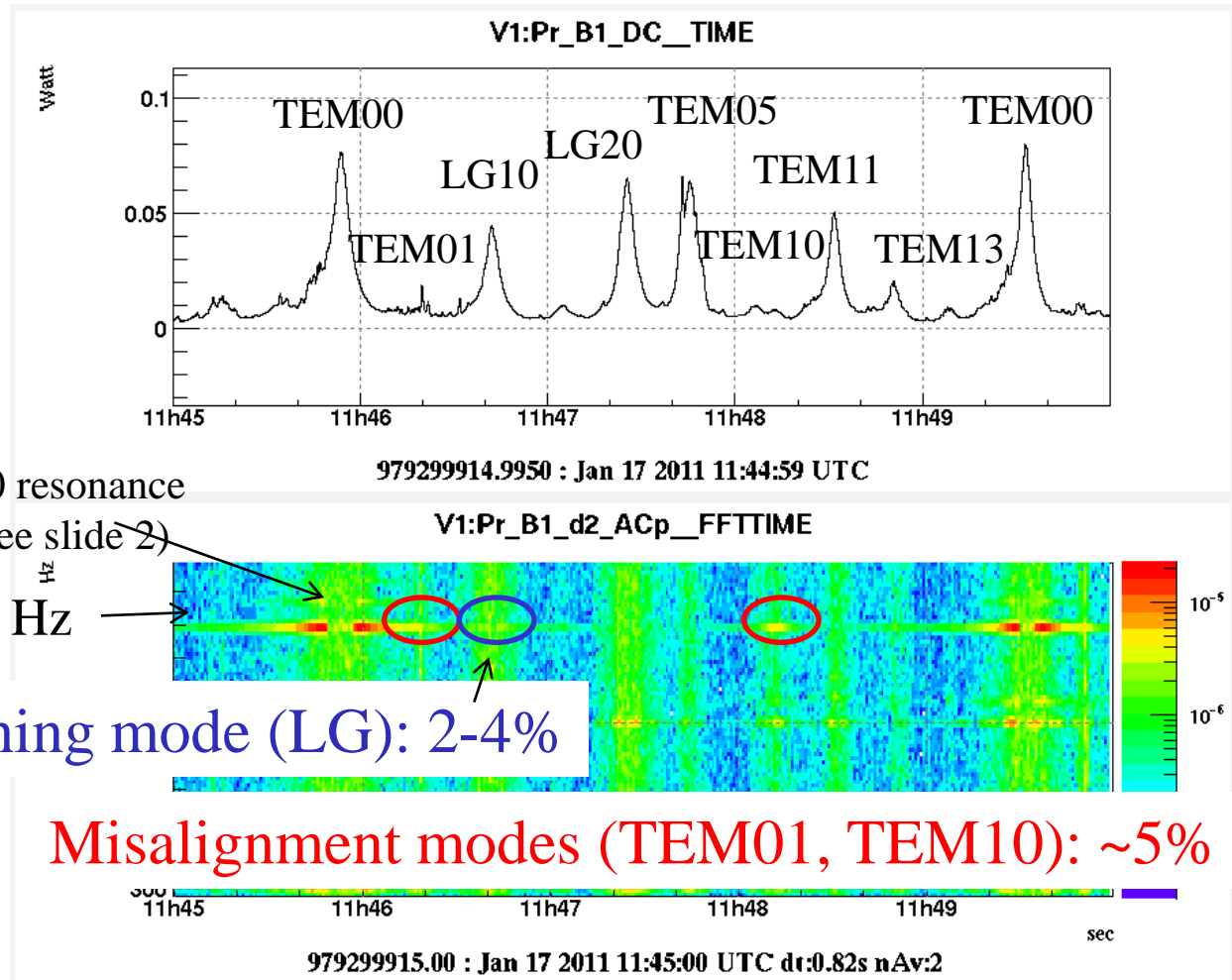
979298505.0000 : Jan 17 2011 11:21:30 UTC

SFP signals

TC
4E

OMC matching / alignment

Measure the fraction of GW signal in each mode: use the differential line at 379Hz



B1_ACp saturates on TEM00 resonance
(due to bad working point, see slide 2)

379 Hz

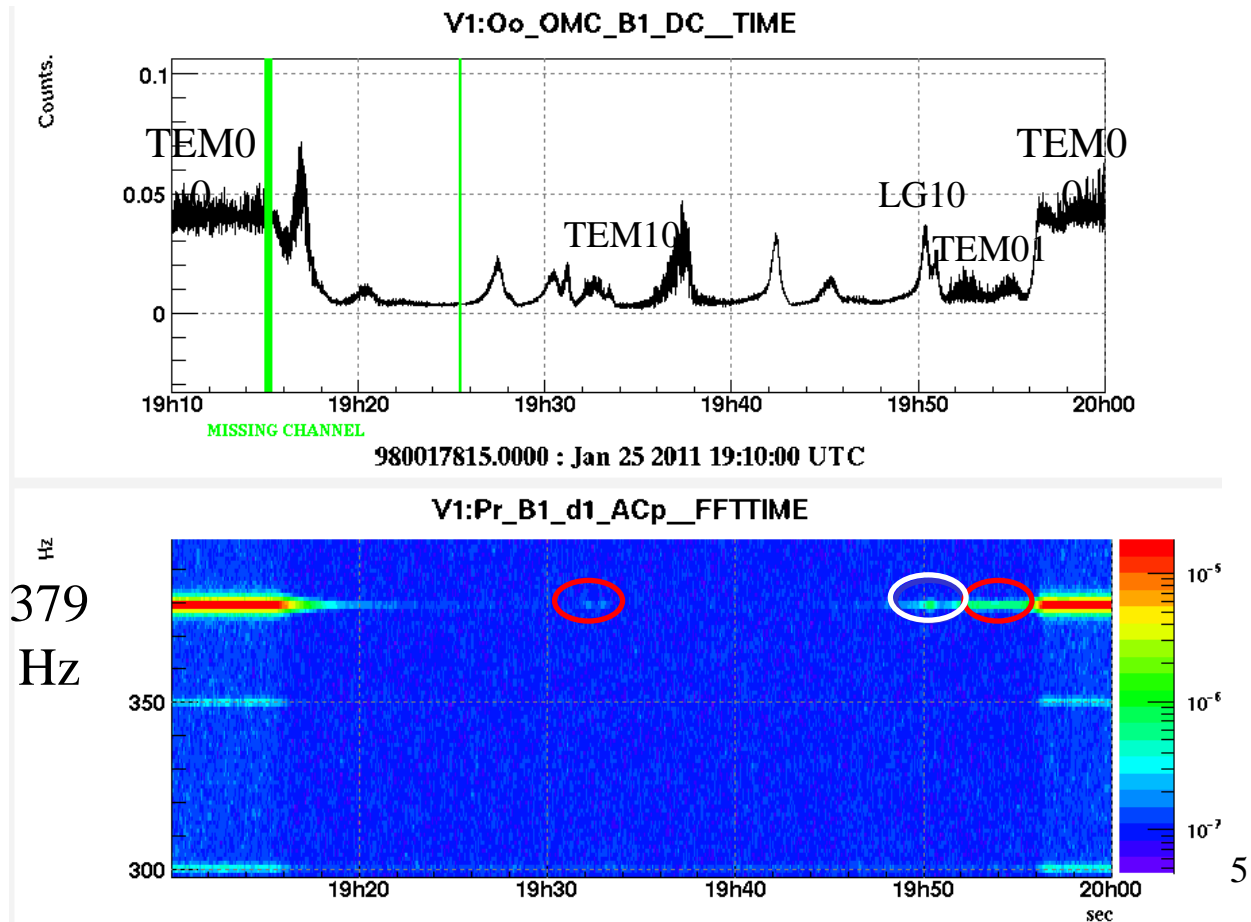
Mismatching mode (LG): 2-4%

Misalignment modes (TEM01, TEM10): ~5%

⇒ Signal loss due to OMC mismatching: ~10%

OMC matching/alignment

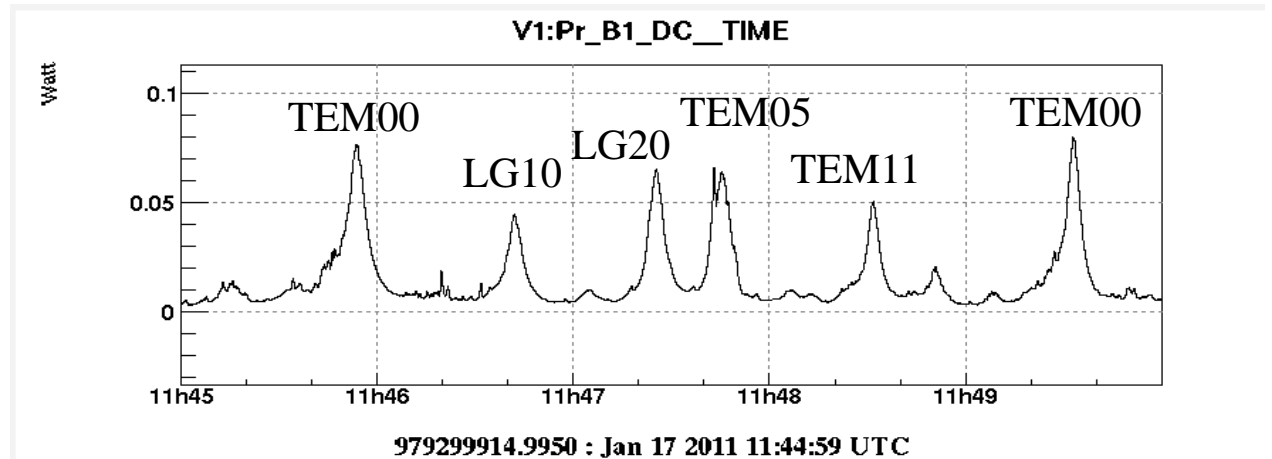
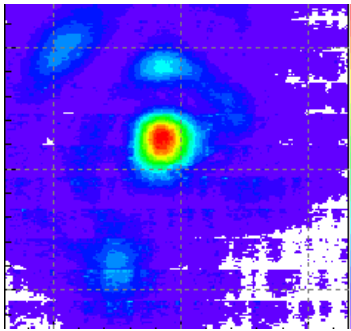
- Same analysis at step8 (same CHRoCC working point):
similar results for matching/alignment
 - ⇒ TEM10 and TEM01 are partly due to OMC misalignment
 - ⇒ other HO mostly reflect dark fringe content



Dark fringe content: HOMs

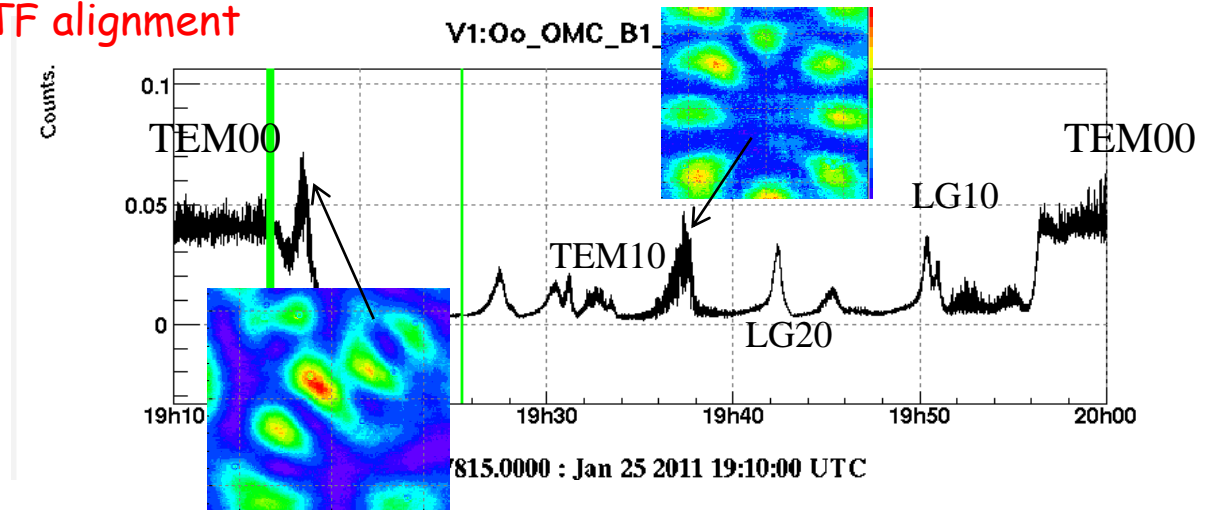
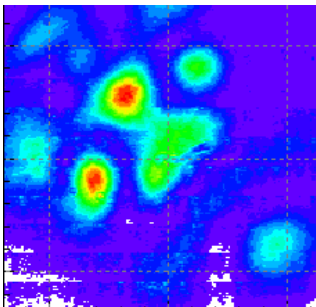
- STEP 12: HO modes mostly due to ROCs and input beam mismatching and astigmatism

B1p at step 12



- STEP 8: HO modes due to poor ITF alignment
 - HO mode very close to TEM00 => OMC lock very difficult
 - => need to improve ITF alignment

B1p at step 8

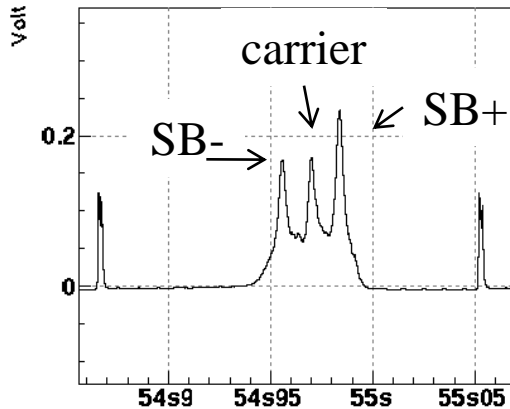


Dark fringe content: carrier/sidebands

- What are the HO modes made of: fraction of carrier and sidebands?
 - Use the scanning Fabry-Perot located after the OMC

OMC on TEM00

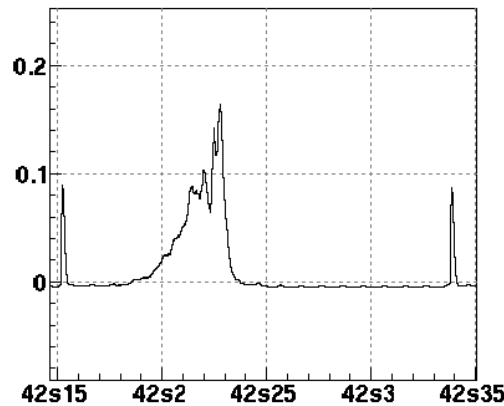
V1:Pr_B1_SFP_TIME



979298649.8558 : Jan 17 2011 11:23:54 UTC

OMC on LG10

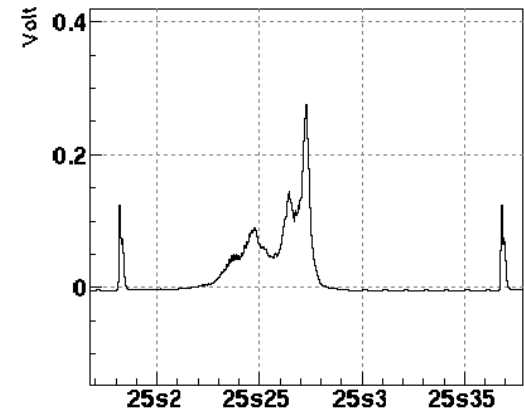
V1:Pr_B1_SFP_TIME



979300017.1460 : Jan 17 2011 11:46:42 UTC

OMC on LG20

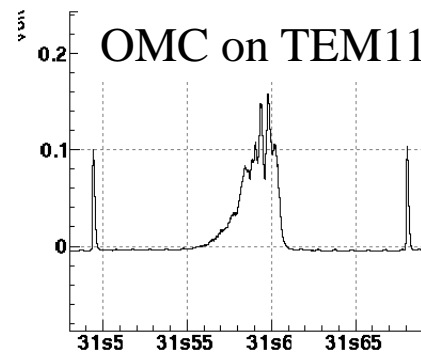
V1:Pr_B1_SFP_TIME



979300060.1670 : Jan 17 2011 11:47:25 UTC

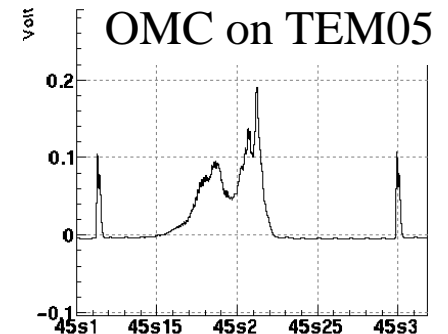
- Content is clear for TEM00:
all 3 components roughly the same
- Decomposition unclear for other modes.
It seems that SB+ dominates (?)
To be investigated

V1:Pr_B1_SFP_TIME



979300126.4801 : Jan 17 2011 11:48:31 UTC

V1:Pr B1 SFP TIME

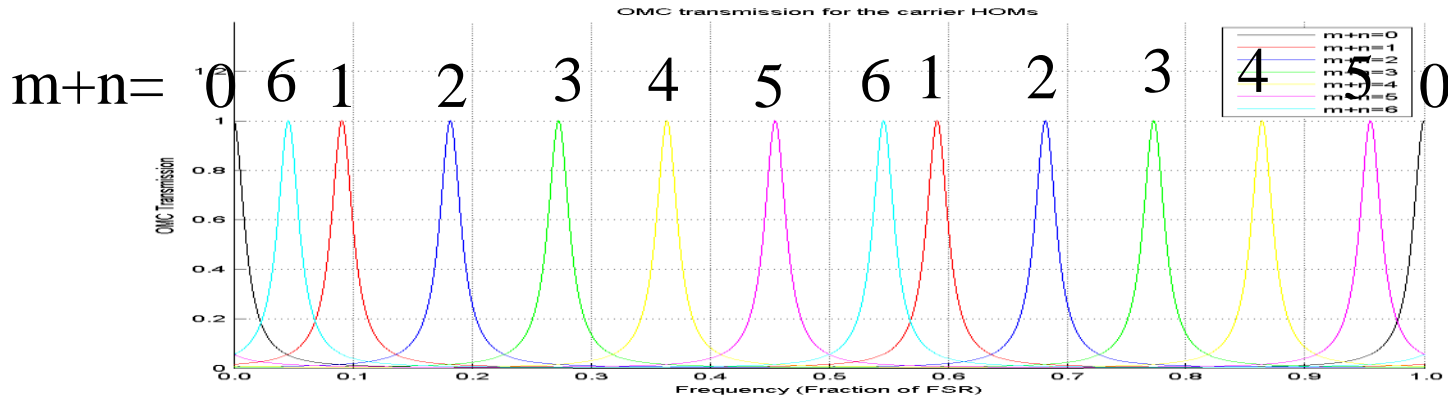


979300080.0998 : Jan 17 2011 11:47:45 UTC

OMC and Fabry-Perot transmission for HOMs

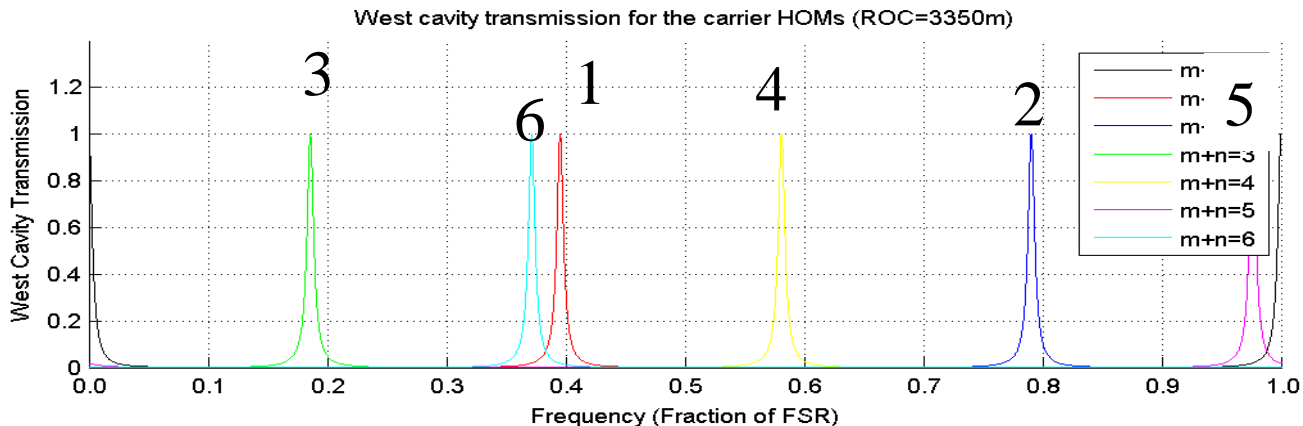
OMC:

- TEM_{nm} ($m+n=5$) close to TEM₀₀ resonance only if n odd ($T \sim 7-8\%$)
- TEM_{nm} ($m+n=11$) resonates with TEM₀₀ ($T \sim 100\%$)



Fabry-Perot cavities: ROC=3350m

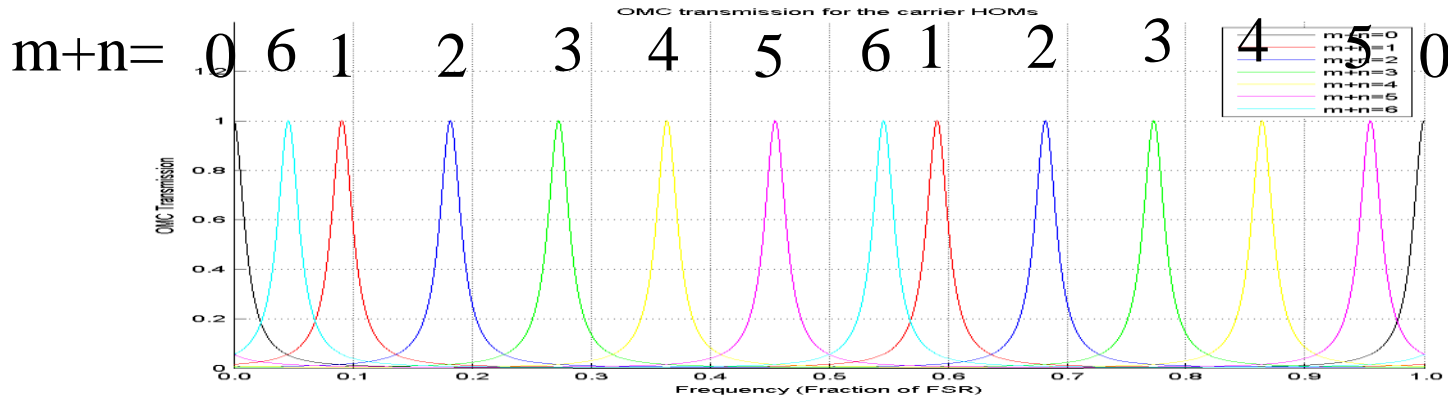
$m+n=$



OMC and Fabry-Perot transmission for HOMs

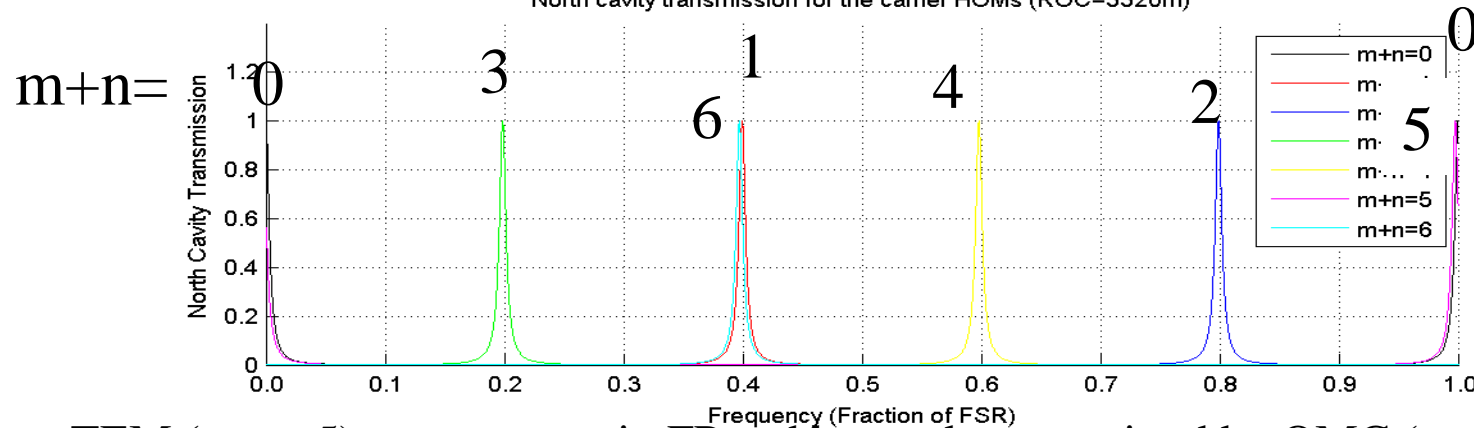
OMC:

- TEM_{nm} ($m+n=5$) close to TEM₀₀ resonance only if n odd ($T \sim 7-8\%$)
- TEM_{nm} ($m+n=11$) resonates with TEM₀₀ ($T \sim 100\%$)



Fabry-Perot cavities: ROC=3320m

North cavity transmission for the carrier HOMs (ROC=3320m)



\Rightarrow TEM ($m+n=5$) can resonate in FP and be partly transmitted by OMC (n odd only) if end mirror ROC = 3320 m

This has not yet been observed but better to avoid ROC=3320m

Measurement of FP losses on TEM00

The asymmetry of losses (on TEM00) inside the Fabry-Perot cavities drive:

- the contrast defect
- the coupling of frequency noise (and potentially other laser noises)

⇒ Impact on Virgo+ sensitivity

The Fabry-Perot losses can be measured by comparing the power reflected by the FP cavity when it is ON or OFF resonance:

- was done in November (logbook 28354, Vajente, Mantovani, Marque) using B1p
=> measure total FP losses (all modes): NA=130ppm, WA=580ppm
- what is important (contrast defect, coupling of frequency noise) is the losses for the TEM00
⇒ Need to use B1 with OMC locked on TEM00

We should redo this measurement using B1

Conclusions

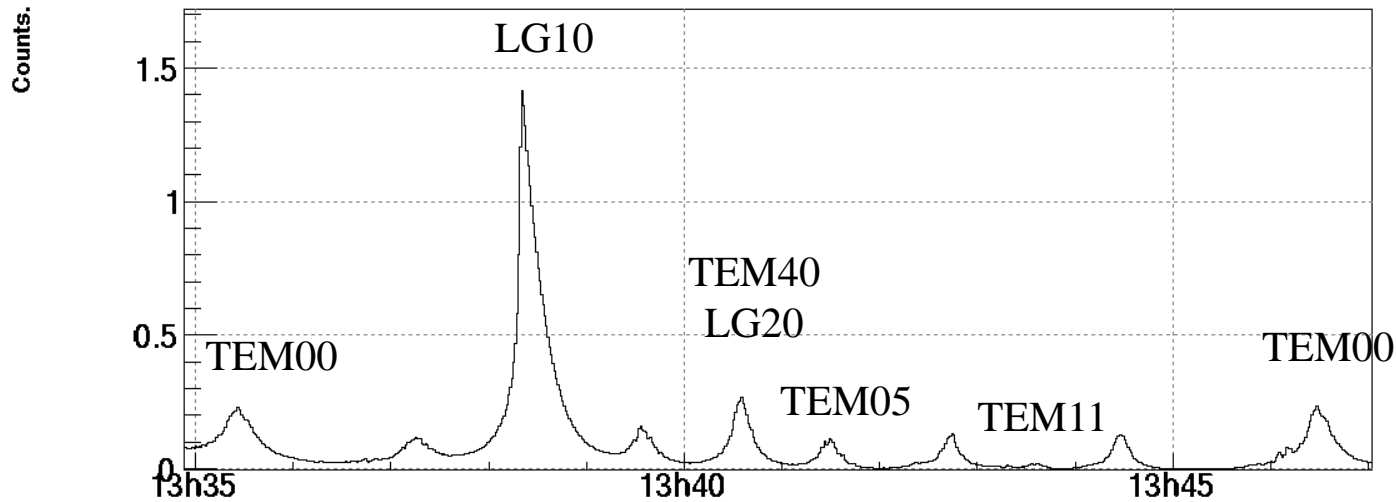
- Large offset on DARM working point during OMC scans
 - ⇒ Analysis of OMC scans more difficult / less precise
 - ⇒ Need to subtract this offset for future scans
- OMC lock difficult at step 8 due to HO modes from ITF misalignment
 - ⇒ Need to improve the ITF alignment at step 8
- Loss of GW signal due to OMC mismatching/misalignment: ~10%
- OMC scans can be used to extract the dark fringe content in HOMs
 - Fraction of carrier/ sidebands is unclear for HOMs (strange SFP signals)
This information would be useful to understand what is really due to ROCs and input beam mismatching
- Other measurement to be done with the OMC: FP losses on TEM00

OMC scans without/with CHRoCC at 300 degrees

dataDisplay v9r12p9 : started by tournefi on Feb 7 2011 09:06:58 UTC

No CHRoCC

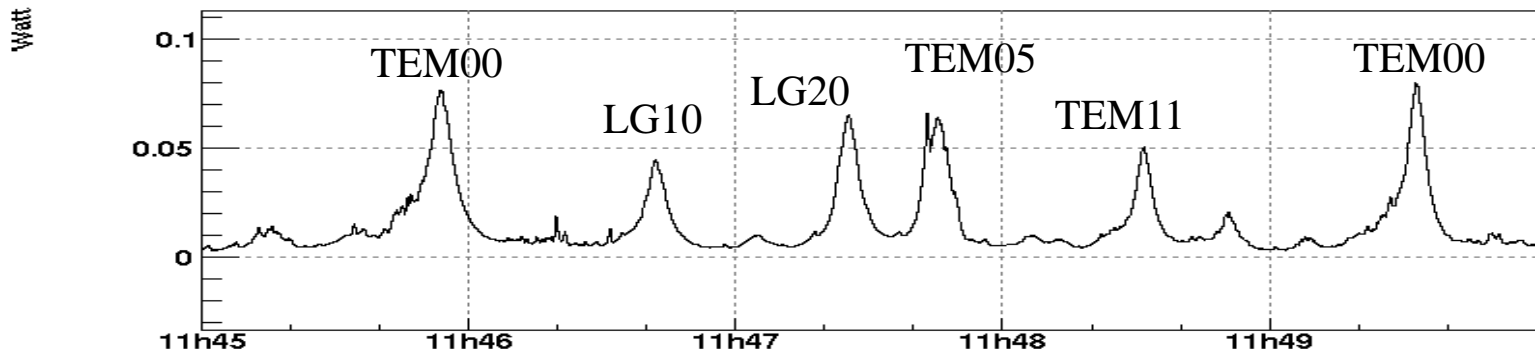
V1:Oo_OMC_B1_DC_mean__TIME



972308108.0000 : Oct 28 2010 13:34:53 UTC

CHRoCC at 310 degrees

V1:Pr_B1_DC__TIME



979299914.9950 : Jan 17 2011 11:44:59 UTC